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Space-time		Line-element	Chapter
Eddington-Finkelstein coordinates	ds^2	$= -\left(1 - \frac{2m}{r}\right) dv^2 + 2dvdr + r^2(d\theta^2 + \sin^2 \theta d\phi^2)$	11
Kerr Spinning black hole	ds^2	$= \left(1 - \frac{2mr}{\Sigma}\right) dt^2 + \frac{4amr \sin^2 \theta}{\Sigma} dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 - \left(r^2 + a^2 + \frac{2a^2mr \sin^2 \theta}{\Sigma}\right) \sin^2 \theta d\phi^2$	2, 11
Kruskal coordinates	ds^2	$= \frac{32m^3}{r} e^{-\frac{r}{2m}} (dv^2 - du^2) - r^2(d\theta + \sin^2 \theta d\phi)$	11
Schwarzschild metric	ds^2	$= -\left(1 - \frac{2m}{r}\right) dt^2 + \left(1 - \frac{2m}{r}\right)^{-1} dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2$	10, 11, 12
Schwarzschild metric: general solution	ds^2	$= e^{2\nu(r)} dt^2 - e^{2\lambda(r)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$	11
Schwarzschild metric: general time dependent	ds^2	$= e^{2\nu(t,r)} dt^2 - e^{2\lambda(t,r)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$	11
Schwarzschild metric: The effect of the cosmological constant over the scale of the solar system	ds^2	$= \left(1 + \frac{1}{3}\Lambda r^2\right) dt^2 - \frac{dr^2}{1 + \frac{1}{3}\Lambda r^2}$	11
Schwarzschild metric: The general Schwarzschild metric with nonzero cosmological constant.	ds^2	$= f(r) dt^2 - \frac{1}{f(r)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$	11
Schwarzschild metric: $\theta = \frac{\pi}{2}$	ds^2	$= -\left(1 - \frac{2m}{r}\right) dt^2 + \left(1 - \frac{2m}{r}\right)^{-1} dr^2 + r^2 d\phi^2$	11

11 The Schwarzschild Spacetime

11.1 ^aThe general Schwarzschild metric

The line element

$$ds^2 = e^{2\nu(r)} dt^2 - e^{2\lambda(r)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

11.1.1 ^bGeodesic equations and Christoffel symbols.

To find the geodesic we use

$$0 = \frac{d}{ds} \left(\frac{\partial K}{\partial \dot{x}^a} \right) - \frac{\partial K}{\partial x^a}$$

$$K = \frac{1}{2} g_{ab} \dot{x}^a \dot{x}^b$$

$$= \frac{1}{2} e^{2\nu(r)} \dot{t}^2 - \frac{1}{2} e^{2\lambda(r)} \dot{r}^2 - \frac{1}{2} r^2 \dot{\theta}^2 - \frac{1}{2} r^2 \sin^2 \theta \dot{\phi}^2$$

$x^a = t$:

$$\frac{\partial K}{\partial t} = 0$$

$$\frac{\partial K}{\partial \dot{t}} = e^{2\nu(r)} \dot{t}$$

$$\frac{d}{ds} \left(\frac{\partial K}{\partial \dot{t}} \right) = 2e^{2\nu(r)} \frac{d\nu}{dr} \dot{r} \dot{t} + e^{2\nu(r)} \ddot{t}$$

$$\Rightarrow 0 = 2e^{2\nu(r)} \frac{d\nu}{dr} \dot{r} \dot{t} + e^{2\nu(r)} \ddot{t}$$

$$\Leftrightarrow 0 = \ddot{t} + 2 \frac{d\nu}{dr} \dot{r} \dot{t}$$

$x^a = r$:

$$\frac{\partial K}{\partial r} = e^{2\nu(r)} \frac{d\nu}{dr} \dot{t}^2 - e^{2\lambda(r)} \frac{d\lambda}{dr} \dot{r}^2 - r \dot{\theta}^2 - r \sin^2 \theta \dot{\phi}^2$$

$$\frac{\partial K}{\partial \dot{r}} = -e^{2\lambda(r)} \dot{r}$$

$$\frac{d}{ds} \left(\frac{\partial K}{\partial \dot{r}} \right) = -2e^{2\lambda(r)} \frac{d\lambda}{dr} \dot{r}^2 - e^{2\lambda(r)} \ddot{r}$$

$$\Rightarrow 0 = -e^{2\lambda(r)} \frac{d\lambda}{dr} \dot{r}^2 - e^{2\lambda(r)} \ddot{r} - e^{2\nu(r)} \frac{d\nu}{dr} \dot{t}^2 + r \dot{\theta}^2 + r \sin^2 \theta \dot{\phi}^2$$

$$\Leftrightarrow 0 = \ddot{r} + \frac{d\lambda}{dr} \dot{r}^2 + e^{2(\nu(r)-\lambda(r))} \frac{d\nu}{dr} \dot{t}^2 - r e^{-2\lambda(r)} \dot{\theta}^2 - r \sin^2 \theta e^{-2\lambda(r)} \dot{\phi}^2$$

$x^a = \theta$:

$$\frac{\partial K}{\partial \theta} = -r^2 \cos \theta \sin \theta \dot{\phi}^2$$

$$\frac{\partial K}{\partial \dot{\theta}} = -r^2 \dot{\theta}$$

$$\frac{d}{ds} \left(\frac{\partial K}{\partial \dot{\theta}} \right) = -2r \dot{r} \dot{\theta} - r^2 \ddot{\theta}$$

$$\Rightarrow 0 = -2r \dot{r} \dot{\theta} - r^2 \ddot{\theta} + r^2 \cos \theta \sin \theta \dot{\phi}^2$$

$$\Leftrightarrow 0 = \ddot{\theta} + \frac{2}{r} \dot{r} \dot{\theta} - \cos \theta \sin \theta \dot{\phi}^2$$

$x^a = \phi$:

$$\frac{\partial K}{\partial \phi} = 0$$

$$\frac{\partial K}{\partial \dot{\phi}} = -r^2 \sin^2 \theta \dot{\phi}$$

$$\frac{d}{ds} \left(\frac{\partial K}{\partial \dot{\phi}} \right) = -2r \sin^2 \theta \dot{r} \dot{\phi} - 2r^2 \cos \theta \sin \theta \dot{\theta} \dot{\phi} - 2r^2 \sin^2 \theta \ddot{\phi}$$

$$\Rightarrow 0 = -2r \sin^2 \theta \dot{r} \dot{\phi} - 2r^2 \cos \theta \sin \theta \dot{\theta} \dot{\phi} - r^2 \sin^2 \theta \ddot{\phi}$$

$$\Leftrightarrow 0 = \ddot{\phi} + \frac{2}{r} \dot{r} \dot{\phi} + 2 \cot \theta \dot{\theta} \dot{\phi}$$

11.1.1.1 Collecting the results

$$0 = \ddot{t} + 2 \frac{d\nu}{dr} \dot{r} \dot{t}$$

$$0 = \ddot{r} + \frac{d\lambda}{dr} \dot{r}^2 + e^{2(\nu(r)-\lambda(r))} \frac{d\nu}{dr} \dot{t}^2 - r e^{-2\lambda(r)} \dot{\theta}^2 - r \sin^2 \theta e^{-2\lambda(r)} \dot{\phi}^2$$

$$0 = \ddot{\theta} + \frac{2}{r} \dot{r} \dot{\theta} - \cos \theta \sin \theta \dot{\phi}^2$$

$$0 = \ddot{\phi} + \frac{2}{r} \dot{r} \dot{\phi} + 2 \cot \theta \dot{\theta} \dot{\phi}$$

11.1.2 The Christoffel symbols

Now we can find the Christoffel symbols from the equation

$$0 = \frac{d^2 x^a}{ds^2} + \Gamma^a_{bc} \frac{dx^b}{ds} \frac{dx^c}{ds}$$

11.1.2.1 Collecting the results

$$\begin{aligned} \Gamma^t_{rt} &= \Gamma^t_{tr} = \frac{dv}{dr} & \Gamma^\theta_{r\theta} &= \Gamma^\theta_{\theta r} = \frac{1}{r} \\ \Gamma^r_{rr} &= \frac{d\lambda}{dr} & \Gamma^\theta_{\phi\phi} &= -\cos \theta \sin \theta \\ \Gamma^r_{tt} &= e^{2(v(r)-\lambda(r))} \frac{dv}{dr} & \Gamma^\phi_{r\phi} &= \Gamma^\phi_{\phi r} = \frac{1}{r} \\ \Gamma^r_{\theta\theta} &= -r e^{-2\lambda(r)} & \Gamma^\phi_{\theta\phi} &= \Gamma^\phi_{\phi\theta} = \cot \theta \\ \Gamma^r_{\phi\phi} &= -r \sin^2 \theta e^{-2\lambda(r)} \end{aligned}$$

11.1.3 The Riemann and Ricci tensor of the general Schwarzschild metric

The line element:

$$ds^2 = e^{2v(r)} dt^2 - e^{2\lambda(r)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

The Basis one forms

$$\begin{aligned} \omega^{\hat{t}} &= e^{v(r)} dt & dt &= e^{-v(r)} \omega^{\hat{t}} \\ \omega^{\hat{r}} &= e^{\lambda(r)} dr & dr &= e^{-\lambda(r)} \omega^{\hat{r}} \\ \omega^{\hat{\theta}} &= r d\theta & d\theta &= \frac{1}{r} \omega^{\hat{\theta}} \\ \omega^{\hat{\phi}} &= r \sin \theta d\phi & d\phi &= \frac{1}{r \sin \theta} \omega^{\hat{\phi}} \end{aligned}$$

$$\eta^{ij} = \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{pmatrix}$$

Cartan's First Structure equation and the curvature one-forms:

$$\begin{aligned} d\omega^{\hat{a}} &= -\Gamma^{\hat{a}}_{\hat{b}} \omega^{\hat{b}} \\ d\omega^{\hat{t}} &= d(e^{v(r)} dt) = v' e^{v(r)} dr \wedge dt = v' e^{-\lambda(r)} \omega^{\hat{r}} \wedge \omega^{\hat{t}} = -v' e^{-\lambda(r)} \omega^{\hat{t}} \wedge \omega^{\hat{r}} = -\Gamma^{\hat{t}}_{\hat{r}} \omega^{\hat{r}} \wedge \omega^{\hat{t}} \\ d\omega^{\hat{r}} &= d(e^{\lambda(r)} dr) = 0 \\ d\omega^{\hat{\theta}} &= d(r d\theta) = dr \wedge d\theta = \frac{1}{r} e^{-\lambda(r)} \omega^{\hat{r}} \wedge \omega^{\hat{\theta}} = -\frac{1}{r} e^{-\lambda(r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{r}} = -\Gamma^{\hat{\theta}}_{\hat{r}} \omega^{\hat{r}} \wedge \omega^{\hat{\theta}} \\ d\omega^{\hat{\phi}} &= d(r \sin \theta d\phi) = \sin \theta dr \wedge d\phi + r \cos \theta d\theta \wedge d\phi \\ &= \frac{1}{r} e^{-\lambda(r)} \omega^{\hat{r}} \wedge \omega^{\hat{\phi}} + \frac{\cot \theta}{r} \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} = -\frac{1}{r} e^{-\lambda(r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{r}} - \frac{\cot \theta}{r} \omega^{\hat{\phi}} \wedge \omega^{\hat{\theta}} \\ &= -\Gamma^{\hat{\phi}}_{\hat{r}} \omega^{\hat{r}} \wedge \omega^{\hat{\phi}} - \Gamma^{\hat{\phi}}_{\hat{\theta}} \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} \end{aligned}$$

11.1.3.1 Collecting the results

$$\begin{aligned} \Gamma^{\hat{t}}_{\hat{r}} &= v' e^{-\lambda(r)} \omega^{\hat{t}} & \Gamma^{\hat{\phi}}_{\hat{r}} &= \frac{1}{r} e^{-\lambda(r)} \omega^{\hat{\phi}} \\ \Gamma^{\hat{\theta}}_{\hat{r}} &= \frac{1}{r} e^{-\lambda(r)} \omega^{\hat{\theta}} & \Gamma^{\hat{\phi}}_{\hat{\theta}} &= \frac{\cot \theta}{r} \omega^{\hat{\phi}} \end{aligned}$$

Summarized in a matrix - Where \hat{a} refers to column and \hat{b} to row:

$$\Gamma^{\hat{a}}_{\hat{b}} = \begin{pmatrix} 0 & v'e^{-\lambda(r)}\omega^{\hat{t}} & 0 & 0 \\ v'e^{-\lambda(r)}\omega^{\hat{t}} & 0 & \frac{1}{r}e^{-\lambda(r)}\omega^{\hat{\theta}} & \frac{1}{r}e^{-\lambda(r)}\omega^{\hat{\phi}} \\ 0 & -\frac{1}{r}e^{-\lambda(r)}\omega^{\hat{\theta}} & 0 & \frac{\cot\theta}{r}\omega^{\hat{\phi}} \\ 0 & -\frac{1}{r}e^{-\lambda(r)}\omega^{\hat{\phi}} & -\frac{\cot\theta}{r}\omega^{\hat{\phi}} & 0 \end{pmatrix}$$

11.1.4 The Ricci Rotation coefficients

$$\begin{aligned} \Gamma^{\hat{a}}_{\hat{b}\hat{c}} &= \Gamma^{\hat{a}}_{\hat{b}\hat{c}}\omega^{\hat{c}} \\ \Gamma^{\hat{t}}_{\hat{r}\hat{t}} &= v'e^{-\lambda(r)} & \Gamma^{\hat{\theta}}_{\hat{r}\hat{\theta}} &= \frac{1}{r}e^{-\lambda(r)} \\ \Gamma^{\hat{r}}_{\hat{t}\hat{t}} &= v'e^{-\lambda(r)} & \Gamma^{\hat{\theta}}_{\hat{\phi}\hat{\phi}} &= -\frac{\cot\theta}{r} \\ \Gamma^{\hat{r}}_{\hat{\theta}\hat{\theta}} &= -\frac{1}{r}e^{-\lambda(r)} & \Gamma^{\hat{\phi}}_{\hat{r}\hat{\phi}} &= \frac{1}{r}e^{-\lambda(r)} \\ \Gamma^{\hat{r}}_{\hat{\phi}\hat{\phi}} &= -\frac{1}{r}e^{-\lambda(r)} & \Gamma^{\hat{\phi}}_{\hat{\theta}\hat{\phi}} &= \frac{\cot\theta}{r} \end{aligned}$$

11.1.5 The curvature two forms

$$\Omega^{\hat{a}}_{\hat{b}} = d\Gamma^{\hat{a}}_{\hat{b}} + \Gamma^{\hat{a}}_{\hat{c}} \wedge \Gamma^{\hat{c}}_{\hat{b}} = \frac{1}{2}R^{\hat{a}}_{\hat{b}\hat{c}\hat{d}}\omega^{\hat{c}} \wedge \omega^{\hat{d}}$$

We need

$$\begin{aligned} d\Gamma^{\hat{r}}_{\hat{t}} &= d(v'e^{-\lambda(r)}\omega^{\hat{t}}) = d(v'e^{v(r)-\lambda(r)}dt) \\ &= (v'' + v'(v' - \lambda'))e^{v(r)-\lambda(r)}dr \wedge dt = (v'' + v'(v' - \lambda'))e^{-2\lambda(r)}\omega^{\hat{r}} \wedge \omega^{\hat{t}} \\ d\Gamma^{\hat{\theta}}_{\hat{r}} &= d\left(\frac{1}{r}e^{-\lambda(r)}\omega^{\hat{\theta}}\right) = d(e^{-\lambda(r)}d\theta) = -\lambda'e^{-\lambda(r)}dr \wedge d\theta = \frac{\lambda'}{r}e^{-2\lambda(r)}\omega^{\hat{\theta}} \wedge \omega^{\hat{r}} \\ d\Gamma^{\hat{\phi}}_{\hat{r}} &= d\left(\frac{1}{r}e^{-\lambda(r)}\omega^{\hat{\phi}}\right) = d(e^{-\lambda(r)}\sin\theta d\phi) \\ &= -\lambda'e^{-\lambda(r)}\sin\theta dr \wedge d\phi + e^{-\lambda(r)}\cos\theta d\theta \wedge d\phi \\ &= -\frac{\lambda'}{r}e^{-2\lambda(r)}\omega^{\hat{r}} \wedge \omega^{\hat{\phi}} + \frac{1}{r^2}e^{-\lambda(r)}\cot\theta\omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} \\ d\Gamma^{\hat{\phi}}_{\hat{\theta}} &= d\left(\frac{\cot\theta}{r}\omega^{\hat{\phi}}\right) = d(\cos\theta d\phi) = -\sin\theta d\theta \wedge d\phi = -\frac{1}{r^2}\omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} \end{aligned}$$

Summerizing:

$$\begin{aligned} d\Gamma^{\hat{r}}_{\hat{t}} &= (v'' + v'(v' - \lambda'))e^{-2\lambda(r)}\omega^{\hat{r}} \wedge \omega^{\hat{t}} \\ d\Gamma^{\hat{\theta}}_{\hat{r}} &= \frac{\lambda'}{r}e^{-2\lambda(r)}\omega^{\hat{\theta}} \wedge \omega^{\hat{r}} \\ d\Gamma^{\hat{\phi}}_{\hat{r}} &= -\frac{\lambda'}{r}e^{-2\lambda(r)}\omega^{\hat{r}} \wedge \omega^{\hat{\phi}} + \frac{1}{r^2}e^{-\lambda(r)}\cot\theta\omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} \\ d\Gamma^{\hat{\phi}}_{\hat{\theta}} &= -\frac{1}{r^2}\omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} \end{aligned}$$

The curvature two-forms

$$\begin{aligned} \Omega^{\hat{r}}_{\hat{t}} &= d\Gamma^{\hat{r}}_{\hat{t}} + \Gamma^{\hat{r}}_{\hat{\theta}} \wedge \Gamma^{\hat{\theta}}_{\hat{t}} + \Gamma^{\hat{r}}_{\hat{\phi}} \wedge \Gamma^{\hat{\phi}}_{\hat{t}} = (v'' + v'(v' - \lambda'))e^{-2\lambda(r)}\omega^{\hat{r}} \wedge \omega^{\hat{t}} \\ \Omega^{\hat{\theta}}_{\hat{t}} &= d\Gamma^{\hat{\theta}}_{\hat{t}} + \Gamma^{\hat{\theta}}_{\hat{r}} \wedge \Gamma^{\hat{r}}_{\hat{t}} + \Gamma^{\hat{\theta}}_{\hat{\phi}} \wedge \Gamma^{\hat{\phi}}_{\hat{t}} = \frac{1}{r}e^{-\lambda(r)}\omega^{\hat{\theta}} \wedge v'e^{-\lambda(r)}\omega^{\hat{t}} = \frac{v'}{r}e^{-2\lambda(r)}\omega^{\hat{\theta}} \wedge \omega^{\hat{t}} \\ \Omega^{\hat{\phi}}_{\hat{t}} &= d\Gamma^{\hat{\phi}}_{\hat{t}} + \Gamma^{\hat{\phi}}_{\hat{r}} \wedge \Gamma^{\hat{r}}_{\hat{t}} + \Gamma^{\hat{\phi}}_{\hat{\theta}} \wedge \Gamma^{\hat{\theta}}_{\hat{t}} = \frac{v'}{r}e^{-2\lambda(r)}\omega^{\hat{\phi}} \wedge \omega^{\hat{t}} \\ \Omega^{\hat{\theta}}_{\hat{r}} &= d\Gamma^{\hat{\theta}}_{\hat{r}} + \Gamma^{\hat{\theta}}_{\hat{t}} \wedge \Gamma^{\hat{t}}_{\hat{r}} + \Gamma^{\hat{\theta}}_{\hat{\phi}} \wedge \Gamma^{\hat{\phi}}_{\hat{r}} = \frac{\lambda'}{r}e^{-2\lambda(r)}\omega^{\hat{\theta}} \wedge \omega^{\hat{r}} \end{aligned}$$

$$\begin{aligned}
\Omega^{\hat{\phi}}_{\hat{r}} &= d\Gamma^{\hat{\phi}}_{\hat{r}} + \Gamma^{\hat{\phi}}_{\hat{t}} \wedge \Gamma^{\hat{t}}_{\hat{r}} + \Gamma^{\hat{\phi}}_{\hat{\theta}} \wedge \Gamma^{\hat{\theta}}_{\hat{r}} \\
&= -\frac{\lambda'}{r} e^{-2\lambda(r)} \omega^{\hat{r}} \wedge \omega^{\hat{\phi}} + \frac{1}{r^2} e^{-\lambda(r)} \cot \theta \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} + \frac{\cot \theta}{r} \omega^{\hat{\phi}} \wedge \frac{1}{r} e^{-\lambda(r)} \omega^{\hat{\theta}} \\
&= -\frac{\lambda'}{r} e^{-2\lambda(r)} \omega^{\hat{r}} \wedge \omega^{\hat{\phi}} \\
\Omega^{\hat{\phi}}_{\hat{\theta}} &= d\Gamma^{\hat{\phi}}_{\hat{\theta}} + \Gamma^{\hat{\phi}}_{\hat{t}} \wedge \Gamma^{\hat{t}}_{\hat{\theta}} + \Gamma^{\hat{\phi}}_{\hat{r}} \wedge \Gamma^{\hat{r}}_{\hat{\theta}} \\
&= -\frac{1}{r^2} \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} - \frac{1}{r^2} e^{-2\lambda(r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{\theta}} \\
&= \frac{(1 - e^{-2\lambda(r)})}{r^2} \omega^{\hat{\phi}} \wedge \omega^{\hat{\theta}}
\end{aligned}$$

11.1.5.1 Collecting the results

$$\begin{aligned}
\Omega^{\hat{r}}_{\hat{t}} &= (v'' + v'(v' - \lambda')) e^{-2\lambda(r)} \omega^{\hat{r}} \wedge \omega^{\hat{t}} & \Omega^{\hat{\theta}}_{\hat{r}} &= \frac{\lambda'}{r} e^{-2\lambda(r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{r}} \\
\Omega^{\hat{\theta}}_{\hat{t}} &= \frac{v'}{r} e^{-2\lambda(r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{t}} & \Omega^{\hat{\phi}}_{\hat{r}} &= -\frac{\lambda'}{r} e^{-2\lambda(r)} \omega^{\hat{r}} \wedge \omega^{\hat{\phi}} \\
\Omega^{\hat{\phi}}_{\hat{t}} &= \frac{v'}{r} e^{-2\lambda(r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{t}} & \Omega^{\hat{\phi}}_{\hat{\theta}} &= \frac{(1 - e^{-2\lambda(r)})}{r^2} \omega^{\hat{\phi}} \wedge \omega^{\hat{\theta}}
\end{aligned}$$

Summarized in a matrix - Where \hat{a} refers to column and \hat{b} to row

$$\Omega^{\hat{a}}_{\hat{b}} = \left\{ \begin{array}{cccc}
0 & (v'' + v'(v' - \lambda')) e^{-2\lambda(r)} \omega^{\hat{r}} \wedge \omega^{\hat{t}} & \frac{v'}{r} e^{-2\lambda(r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{t}} & \frac{v'}{r} e^{-2\lambda(r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{t}} \\
S & 0 & \frac{\lambda'}{r} e^{-2\lambda(r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{r}} & \frac{\lambda'}{r} e^{-2\lambda(r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{r}} \\
S & AS & 0 & \frac{(1 - e^{-2\lambda(r)})}{r^2} \omega^{\hat{\phi}} \wedge \omega^{\hat{\theta}} \\
S & AS & AS & 0
\end{array} \right\}$$

The independent elements of the Riemann tensor in the non-coordinate basis

$$\begin{aligned}
R^{\hat{r}}_{\hat{t}\hat{r}\hat{t}} &= (v'' + v'(v' - \lambda')) e^{-2\lambda(r)} & R^{\hat{\phi}}_{\hat{t}\hat{\phi}\hat{t}} &= \frac{v'}{r} e^{-2\lambda(r)} \\
R^{\hat{\theta}}_{\hat{t}\hat{\theta}\hat{t}} &= \frac{v'}{r} e^{-2\lambda(r)} & R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{r}} &= \frac{\lambda'}{r} e^{-2\lambda(r)} \\
R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{r}} &= \frac{\lambda'}{r} e^{-2\lambda(r)} & R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{\theta}} &= \frac{(1 - e^{-2\lambda(r)})}{r^2}
\end{aligned}$$

11.1.5.2 The Ricci tensor

$$\begin{aligned}
R_{\hat{a}\hat{b}} &= R^{\hat{c}}_{\hat{a}\hat{c}\hat{b}} \\
R_{\hat{t}\hat{t}} &= R^{\hat{r}}_{\hat{t}\hat{r}\hat{t}} + R^{\hat{\theta}}_{\hat{t}\hat{\theta}\hat{t}} + R^{\hat{\phi}}_{\hat{t}\hat{\phi}\hat{t}} = \left(v'' + v'(v' - \lambda') + 2\frac{v'}{r} \right) e^{-2\lambda(r)} \\
R_{\hat{r}\hat{r}} &= R^{\hat{c}}_{\hat{r}\hat{c}\hat{r}} = R^{\hat{t}}_{\hat{r}\hat{t}\hat{r}} + R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{r}} + R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{r}} \\
&= -R^{\hat{r}}_{\hat{t}\hat{r}\hat{t}} + R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{r}} + R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{r}} = -\left(v'' + v'(v' - \lambda') - 2\frac{\lambda'}{r} \right) e^{-2\lambda(r)} \\
R_{\hat{\theta}\hat{\theta}} &= R^{\hat{c}}_{\hat{\theta}\hat{c}\hat{\theta}} = R^{\hat{t}}_{\hat{\theta}\hat{t}\hat{\theta}} + R^{\hat{r}}_{\hat{\theta}\hat{r}\hat{\theta}} + R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{\theta}} \\
&= -R^{\hat{\theta}}_{\hat{t}\hat{\theta}\hat{t}} + R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{r}} + R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{\theta}} = -\frac{v'}{r} e^{-2\lambda(r)} + \frac{\lambda'}{r} e^{-2\lambda(r)} + \frac{(1 - e^{-2\lambda(r)})}{r^2} \\
R_{\hat{\phi}\hat{\phi}} &= R^{\hat{c}}_{\hat{\phi}\hat{c}\hat{\phi}} = R^{\hat{t}}_{\hat{\phi}\hat{t}\hat{\phi}} + R^{\hat{r}}_{\hat{\phi}\hat{r}\hat{\phi}} + R^{\hat{\theta}}_{\hat{\phi}\hat{\theta}\hat{\phi}} = -R^{\hat{\phi}}_{\hat{t}\hat{\phi}\hat{t}} + R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{r}} + R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{\theta}}
\end{aligned}$$

$$= -\frac{v'}{r}e^{-2\lambda(r)} + \frac{\lambda'}{r}e^{-2\lambda(r)} + \frac{(1 - e^{-2\lambda(r)})}{r^2}$$

The non-diagonal elements:

$$R_{\hat{a}\hat{b}} = R^{\hat{t}}_{\hat{a}\hat{t}\hat{b}} + R^{\hat{r}}_{\hat{a}\hat{r}\hat{b}} + R^{\hat{\theta}}_{\hat{a}\hat{\theta}\hat{b}} + R^{\hat{\phi}}_{\hat{a}\hat{\phi}\hat{b}} = 0$$

11.1.5.3 Collecting the results

$$R_{\hat{t}\hat{t}} = \left(v'' + v'(v' - \lambda') + 2\frac{v'}{r} \right) e^{-2\lambda(r)}$$

$$R_{\hat{r}\hat{r}} = -\left(v'' + v'(v' - \lambda') - 2\frac{\lambda'}{r} \right) e^{-2\lambda(r)}$$

$$R_{\hat{\theta}\hat{\theta}} = -\frac{v'}{r}e^{-2\lambda(r)} + \frac{\lambda'}{r}e^{-2\lambda(r)} + \frac{(1 - e^{-2\lambda(r)})}{r^2}$$

$$R_{\hat{\phi}\hat{\phi}} = -\frac{v'}{r}e^{-2\lambda(r)} + \frac{\lambda'}{r}e^{-2\lambda(r)} + \frac{(1 - e^{-2\lambda(r)})}{r^2}$$

11.1.6 ^deThe Vacuum equations

$$G_{\hat{a}\hat{b}} = R_{\hat{a}\hat{b}} - \frac{1}{2}\eta_{\hat{a}\hat{b}}R = 0$$

$$\Rightarrow^1 R_{\hat{a}\hat{b}} = 0$$

So in order to find the parameters v and λ we can solve:

$$R_{\hat{t}\hat{t}} = \left(v'' + v'(v' - \lambda') + 2\frac{v'}{r} \right) e^{-2\lambda(r)} = 0 \quad (11.1.)$$

$$R_{\hat{r}\hat{r}} = -\left(v'' + v'(v' - \lambda') - 2\frac{\lambda'}{r} \right) e^{-2\lambda(r)} = 0 \quad (11.2.)$$

$$R_{\hat{\theta}\hat{\theta}} = R_{\hat{\phi}\hat{\phi}} = -\frac{v'}{r}e^{-2\lambda(r)} + \frac{\lambda'}{r}e^{-2\lambda(r)} + \frac{(1 - e^{-2\lambda(r)})}{r^2} = 0 \quad (11.3.)$$

Adding eq. (11.1.) and eq. (11.2.)

$$R_{\hat{t}\hat{t}} + R_{\hat{r}\hat{r}} = \frac{2}{r}(v' + \lambda')e^{-2\lambda(r)} = 0$$

$$\Rightarrow v' = -\lambda' \quad (11.4.)$$

Substituting (11.4.) into (11.3.)

$$\Rightarrow 0 = 2\frac{\lambda'}{r}e^{-2\lambda(r)} + \frac{(1 - e^{-2\lambda(r)})}{r^2}$$

$$\Rightarrow 0 = (2r\lambda' - 1)e^{-2\lambda(r)} + 1$$

Defining

$$\gamma = e^{-2\lambda(r)}$$

$$\Rightarrow \lambda' = 2 - \frac{1}{2}\gamma'e^{2\lambda(r)}$$

$$\begin{aligned} \Rightarrow 0 &= (2r\lambda' - 1)e^{-2\lambda(r)} + 1 = 2r\lambda'e^{-2\lambda(r)} - e^{-2\lambda(r)} + 1 \\ &= 2r\left(-\frac{1}{2}\gamma'e^{2\lambda(r)}\right)e^{-2\lambda(r)} - \gamma + 1 = -r\gamma' - \gamma + 1 \end{aligned}$$

$$\Rightarrow \gamma = 31 - \frac{k}{r}$$

¹ See the chapter named: The Vacuum Einstein Equations

² $\gamma' = -2\lambda'e^{-2\lambda(r)} \Rightarrow \lambda' = -\frac{1}{2}\gamma'e^{2\lambda(r)}$

³ $\gamma e^{\int \frac{1}{r} dr} = \int \frac{1}{r} e^{\int \frac{1}{r} dr} dr + k \Rightarrow \gamma e^{\ln(r)} = \int \frac{1}{r} e^{\ln(r)} dr + k \Rightarrow \gamma r = r + k$ (Spiegel, 1990) (18.2)

$$\Rightarrow e^{2\lambda(r)} = \frac{1}{\gamma} = \frac{1}{1 - \frac{k}{r}}$$

$$e^{2\nu(r)} = e^{-2\lambda(r)} = \gamma = 1 - \frac{k}{r}$$

Substituting in the line element

$$ds^2 = e^{2\nu(r)} dt^2 - e^{2\lambda(r)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

$$= \left(1 - \frac{k}{r}\right) dt^2 - \left(\frac{1}{1 - \frac{k}{r}}\right) dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

11.1.7 The meaning of the integration constant: The choice of $k = 2m$

We can use the geodesic equations to justify the choice of $2m$ by investigating the geodesic equations in the classical limit i.e. $r \gg 2m$, $\frac{dt}{d\tau} \rightarrow 1$ and $v = \frac{dr}{d\tau} \ll c$, where v is the velocity and τ is the proper time. We want to investigate the case of a radially infalling particle i.e. $d\theta = 0$ and $d\phi = 0$. We also want to work in SI-units so we have to substitute m by $\frac{Gm}{c^2}$. Also remember that $\dot{t} = \frac{dt}{ds} = \frac{dt}{cd\tau} \rightarrow \frac{1}{c}$, $\dot{r} = \frac{dr}{ds} = \frac{dr}{cd\tau} = \frac{v}{c}$ and $\ddot{r} = \frac{d^2r}{ds^2} = \frac{d^2r}{c^2 d\tau^2} = \frac{a}{c^2}$, where a is the particle acceleration. We use equation (10.38):

$$0 = \ddot{r} - \frac{m}{r(r-2m)} \dot{r}^2 + \frac{m(r-2m)}{r^3} \dot{t}^2 - (r-2m)\dot{\theta}^2 - (r-2m)\sin^2 \theta \dot{\phi}^2$$

Now before we carry on with the physics we also have to be sure that each term in this equation has the same dimension. It turns out that they don't and therefore the third term has to be multiplied by c^2 , in which case each term gets the dimension⁴ $\frac{1}{\text{length}}$.

$$\Rightarrow 0 = \ddot{r} - \frac{m}{r(r-2m)} \dot{r}^2 + \frac{m(r-2m)}{r^3} c^2 \dot{t}^2 - (r-2m)\dot{\theta}^2 - (r-2m)\sin^2 \theta \dot{\phi}^2$$

$d\theta = 0, d\phi = 0$:

$$\Rightarrow 0 = \ddot{r} - \frac{m}{r(r-2m)} \dot{r}^2 + \frac{m(r-2m)}{r^3} c^2 \dot{t}^2$$

$r \gg 2m$:

$$\Rightarrow 0 = \ddot{r} - \frac{m}{r^2} \dot{r}^2 + \frac{m}{r^2} c^2 \dot{t}^2$$

$\dot{t} = \frac{1}{c}, \dot{r} = \frac{v}{c}, \ddot{r} = \frac{a}{c^2}$:

$$\Rightarrow 0 = \frac{a}{c^2} - \frac{m}{r^2} \frac{v^2}{c^2} + \frac{m}{r^2} = \frac{1}{c^2} \left(a - \frac{m}{r^2} v^2 + \frac{mc^2}{r^2} \right)$$

$m \rightarrow \frac{Gm}{c^2}$:

$$\Rightarrow 0 = a - \frac{Gm}{r^2} \left(\frac{v}{c}\right)^2 + \frac{Gm}{r^2}$$

$v \ll c$:

$$\Rightarrow a = -\frac{Gm}{r^2}$$

Multiplying with M on both sides we get precisely the Newtonian gravitational law

$$\Rightarrow F = Ma = -\frac{GMm}{r^2}$$

⁴ This actually originates from the line element of the Schwarzschild metric itself, because in order to get the same dimension of each term, the first term has to be multiplied by c^2 : $ds^2 = \left(1 - \frac{2m}{r}\right) c^2 dt^2 - \left(1 - \frac{2m}{r}\right)^{-1} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$

11.1.8 ^gThe meaning of the coordinate r .

The coordinate r is not a distance from any centre, rather it is related to the area A of a two-dimensional sphere for some fixed t and r .

$$r = \left(\frac{A}{4\pi}\right)^{\frac{1}{2}}$$

How: If you look at the Schwarzschild line-element

$$ds^2 = -\left(1 - \frac{2m}{r}\right) dt^2 + \left(1 - \frac{2m}{r}\right)^{-1} dr^2 + r^2(d\theta^2 + \sin^2 \theta d\phi^2)$$

You can see that embedded in this is the geometry of a two-dimensional sphere with line-element

$$d\Sigma^2 = r^2(d\theta^2 + \sin^2 \theta d\phi^2)$$

The areal element

$$dA = dl^2 dl^3 = r^2 \sin \theta d\theta d\phi$$

$$\Rightarrow A = \int_0^\pi r^2 \sin \theta d\theta \int_0^{2\pi} d\phi = 4\pi r^2$$

11.2 The Schwarzschild geometry

The line element

$$ds^2 = -\left(1 - \frac{2m}{r}\right) dt^2 + \left(1 - \frac{2m}{r}\right)^{-1} dr^2 + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2$$

11.2.1 ^hLength and volume of the Schwarzschild geometry

The spatial part of the Schwarzschild line element is

$$dS^2 = -\left(1 - \frac{2m}{r}\right)^{-1} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

(a) The radial distance between the sphere $r = 2M$ and the sphere $r = 3M$

$$\begin{aligned} l &= \int dl^1 = \int_{2M}^{3M} \frac{1}{\sqrt{1 - \frac{2M}{r}}} dr = \int_{2M}^{3M} \sqrt{\frac{r}{r - 2M}} dr \\ &= {}^5 \left[\sqrt{(r - 2M)r} \right]_{2M}^{3M} + M \int_{2M}^{3M} \frac{dr}{\sqrt{(r - 2M)r}} \\ &= {}^6 \left[\sqrt{(r - 2M)r} + 2M \ln(\sqrt{r} + \sqrt{r - 2M}) \right]_{2M}^{3M} \\ &= \sqrt{M * 3M} + 2M \ln(\sqrt{3M} + \sqrt{M}) - 2M \ln(\sqrt{2M}) = \left(\frac{\sqrt{3}}{2} + \ln\left(\frac{\sqrt{3} + 1}{\sqrt{2}}\right) \right) * 2M \\ &\approx 1,52 * 2M \end{aligned}$$

(b) The spatial volume between the sphere $r = 2M$ and the sphere $r = 3M$

$$\begin{aligned} \mathcal{V} &= \int_{2M}^{3M} dl^1 dl^2 dl^3 = \int_{2M}^{3M} \frac{r^2}{\sqrt{1 - \frac{2M}{r}}} dr \int_0^\pi \sin \theta d\theta \int_0^{2\pi} d\phi = 4\pi * 17,15 * M^3 \\ &= 6,43 * \frac{4\pi}{3} (2M)^3 \end{aligned}$$

11.2.2 ⁱGeodesics in the Schwarzschild Spacetime

To find the geodesic we use the Euler-Lagrange equation

$${}^5 \int \frac{px+q}{\sqrt{ax+b}} dx = \frac{\sqrt{(ax+b)(px+q)}}{a} + \frac{aq-bp}{2a} \int \frac{dx}{\sqrt{(ax+b)(px+q)}} \quad (\text{Spiegel, 1990}) \quad (14.123)$$

$${}^6 \int \frac{dx}{\sqrt{(ax+b)(px+q)}} = \frac{2}{\sqrt{ap}} \ln\left(\sqrt{a(px+q)} + \sqrt{p(ax+b)}\right) \quad (\text{Spiegel, 1990}) \quad (14.120)$$

$$0 = \frac{d}{ds} \left(\frac{\partial F}{\partial \dot{x}^a} \right) - \frac{\partial F}{\partial x^a}$$

where

$$F = \left(1 - \frac{2m}{r}\right) \dot{t}^2 - \left(1 - \frac{2m}{r}\right)^{-1} \dot{r}^2 - r^2 \dot{\theta}^2 - r^2 \sin^2 \theta \dot{\phi}^2$$

$x^a = t$:

$$\frac{\partial F}{\partial t} = 0$$

$$\frac{\partial F}{\partial \dot{t}} = 2 \left(1 - \frac{2m}{r}\right) \dot{t}$$

$$\frac{d}{ds} \left(\frac{\partial F}{\partial \dot{t}} \right) = \frac{4m}{r^2} \dot{r} \dot{t} + 2 \left(1 - \frac{2m}{r}\right) \ddot{t}$$

$$\Rightarrow 0 = \frac{4m}{r^2} \dot{r} \dot{t} + 2 \left(1 - \frac{2m}{r}\right) \ddot{t}$$

$$\Leftrightarrow 0 = \ddot{t} + \frac{2m}{r(r-2m)} \dot{r} \dot{t}$$

$x^a = r$:

$$\frac{\partial F}{\partial r} = \frac{2m}{r^2} \dot{t}^2 + \frac{2m}{(r-2m)^2} \dot{r}^2 - 2r \dot{\theta}^2 - 2r \sin^2 \theta \dot{\phi}^2$$

$$\frac{\partial F}{\partial \dot{r}} = -2 \left(1 - \frac{2m}{r}\right)^{-1} \dot{r}$$

$$\frac{d}{ds} \left(\frac{\partial F}{\partial \dot{r}} \right) = -2 \left(1 - \frac{2m}{r}\right)^{-1} \ddot{r} + 2 \left(1 - \frac{2m}{r}\right)^{-2} \frac{2m}{r^2} \dot{r}^2 = -2 \left(1 - \frac{2m}{r}\right)^{-1} \ddot{r} + \frac{4m}{(r-2m)^2} \dot{r}^2$$

$$\Rightarrow 0 = -2 \left(1 - \frac{2m}{r}\right)^{-1} \ddot{r} + \frac{2m}{(r-2m)^2} \dot{r}^2 - \frac{2m}{r^2} \dot{t}^2 + 2r \dot{\theta}^2 + 2r \sin^2 \theta \dot{\phi}^2$$

$$\Leftrightarrow 0 = \ddot{r} - \frac{m}{r(r-2m)} \dot{r}^2 + \frac{m(r-2m)}{r^3} \dot{t}^2 - (r-2m) \dot{\theta}^2 - (r-2m) \sin^2 \theta \dot{\phi}^2$$

$x^a = \theta$:

$$\frac{\partial F}{\partial \theta} = -2r^2 \cos \theta \sin \theta \dot{\phi}^2$$

$$\frac{\partial F}{\partial \dot{\theta}} = -2r^2 \dot{\theta}$$

$$\frac{d}{ds} \left(\frac{\partial F}{\partial \dot{\theta}} \right) = -4r \dot{r} \dot{\theta} - 2r^2 \ddot{\theta}$$

$$\Rightarrow 0 = -4r \dot{r} \dot{\theta} - 2r^2 \ddot{\theta} + 2r^2 \cos \theta \sin \theta \dot{\phi}^2$$

$$\Leftrightarrow 0 = \ddot{\theta} + \frac{2}{r} \dot{r} \dot{\theta} - \cos \theta \sin \theta \dot{\phi}^2$$

$x^a = \phi$:

$$\frac{\partial F}{\partial \phi} = 0$$

$$\frac{\partial F}{\partial \dot{\phi}} = -2r^2 \sin^2 \theta \dot{\phi}$$

$$\frac{d}{ds} \left(\frac{\partial F}{\partial \dot{\phi}} \right) = -4r \sin^2 \theta \dot{r} \dot{\phi} - 4r^2 \cos \theta \sin \theta \dot{\theta} \dot{\phi} - 2r^2 \sin^2 \theta \ddot{\phi}$$

$$\Rightarrow 0 = -4r \sin^2 \theta \dot{r} \dot{\phi} - 4r^2 \cos \theta \sin \theta \dot{\theta} \dot{\phi} - 2r^2 \sin^2 \theta \ddot{\phi}$$

$$\Leftrightarrow 0 = \ddot{\phi} + \frac{2}{r} \dot{r} \dot{\phi} + 2 \cot \theta \dot{\theta} \dot{\phi}$$

11.2.2.1 Collecting the results

$$0 = \ddot{t} + \frac{2m}{r(r-2m)} \dot{r} \dot{t}$$

$$0 = \ddot{r} - \frac{m}{r(r-2m)} \dot{r}^2 + \frac{m(r-2m)}{r^3} \dot{t}^2 - (r-2m)\dot{\theta}^2 - (r-2m)\sin^2\theta \dot{\phi}^2$$

$$0 = \ddot{\theta} + \frac{2}{r} \dot{r} \dot{\theta} - \cos\theta \sin\theta \dot{\phi}^2$$

$$0 = \ddot{\phi} + \frac{2}{r} \dot{r} \dot{\phi} + 2 \cot\theta \dot{\theta} \dot{\phi}$$

11.2.3 The Riemann tensor of the Schwarzschild metric

Solving the vacuum equations we find $\nu = -\lambda$ and $e^{2\nu} = 1 - \frac{2m}{r}$, $e^{2\lambda} = \left(1 - \frac{2m}{r}\right)^{-1}$

The line element:

$$ds^2 = \left(1 - \frac{2m}{r}\right) dt^2 - \left(1 - \frac{2m}{r}\right)^{-1} dr^2 - r^2 d\theta^2 - r^2 \sin^2\theta d\phi^2$$

Now we can find the independent elements of the Riemann tensor in the non-coordinate basis:

$$\nu' = \frac{1}{2} e^{-2\nu} \frac{d}{dr} (e^{2\nu}) = \frac{1}{2} e^{-2\nu} \frac{d}{dr} \left(1 - \frac{2m}{r}\right) = \frac{1}{2} e^{-2\nu} \left(\frac{2m}{r^2}\right) = \frac{m}{r^2 - r2m}$$

$$\nu'' = \frac{d}{dr} \left(\frac{1}{2} e^{-2\nu} \frac{d}{dr} (e^{2\nu})\right) = -\nu' e^{-2\nu} \frac{d}{dr} (e^{2\nu}) + \frac{1}{2} e^{-2\nu} \frac{d^2}{dr^2} (e^{2\nu})$$

$$= -2\nu'\nu' + \frac{1}{2} e^{-2\nu} \frac{d}{dr} \left(\frac{2m}{r^2}\right) = -2\nu'\nu' - \frac{2m}{r^3} e^{-2\nu}$$

$$\Rightarrow R^{\hat{r}}_{\hat{t}\hat{r}\hat{t}} = (\nu'' + \nu'(\nu' - \lambda')) e^{-2\lambda(r)} = (\nu'' + 2\nu'\nu') e^{2\nu(r)}$$

$$= \left(-2\nu'\nu' - \frac{2m}{r^3} e^{-2\nu} + 2\nu'\nu'\right) e^{2\nu(r)} = -\frac{2m}{r^3}$$

$$R^{\hat{\theta}}_{\hat{t}\hat{\theta}\hat{t}} = \frac{\nu'}{r} e^{-2\lambda(r)} = \frac{\nu'}{r} e^{2\nu(r)} = \frac{1}{r} \frac{1}{2} e^{-2\nu} \left(\frac{2m}{r^2}\right) e^{2\nu(r)} = \frac{m}{r^3}$$

$$R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{r}} = \frac{\lambda'}{r} e^{-2\lambda(r)} = -\frac{m}{r^3}$$

$$R^{\hat{\phi}}_{\hat{t}\hat{\phi}\hat{t}} = \frac{\nu'}{r} e^{-2\lambda(r)} = \frac{m}{r^3}$$

$$R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{r}} = \frac{\lambda'}{r} e^{-2\lambda(r)} = -\frac{m}{r^3}$$

$$R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{\theta}} = \frac{(1 - e^{-2\lambda(r)})}{r^2} = \frac{(1 - e^{2\nu(r)})}{r^2} = \frac{\left(1 - \left(1 - \frac{2m}{r}\right)\right)}{r^2} = \frac{2m}{r^3}$$

11.2.3.1 Collecting the results

$$R^{\hat{r}}_{\hat{t}\hat{r}\hat{t}} = -\frac{2m}{r^3}$$

$$R^{\hat{\theta}}_{\hat{t}\hat{\theta}\hat{t}} = R^{\hat{\phi}}_{\hat{t}\hat{\phi}\hat{t}} = \frac{m}{r^3}$$

$$R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{r}} = R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{r}} = -\frac{m}{r^3}$$

$$R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{\theta}} = \frac{2m}{r^3}$$

11.3 The general time-dependent Schwarzschild space-time

11.3.1 The Ricci tensor, Ricci scalar and Einstein tensor

The line element:

$$ds^2 = e^{2\nu(t,r)} dt^2 - e^{2\lambda(t,r)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

The Basis one forms

$$\begin{aligned}\omega^{\hat{t}} &= e^{\nu(t,r)} dt & dt &= e^{-\nu(t,r)} \omega^{\hat{t}} \\ \omega^{\hat{r}} &= e^{\lambda(t,r)} dr & dr &= e^{-\lambda(t,r)} \omega^{\hat{r}} \\ \omega^{\hat{\theta}} &= r d\theta & d\theta &= \frac{1}{r} \omega^{\hat{\theta}} \\ \omega^{\hat{\phi}} &= r \sin \theta d\phi & d\phi &= \frac{1}{r \sin \theta} \omega^{\hat{\phi}}\end{aligned}$$

$$\eta^{ij} = \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{pmatrix}$$

Cartan's First Structure equation and the calculation of the Cartan structure coefficients $\Gamma_{\hat{b}\hat{a}}$:

$$\begin{aligned}d\omega^{\hat{a}} &= -\Gamma_{\hat{b}\hat{a}} \wedge \omega^{\hat{b}} \\ d\omega^{\hat{t}} &= d(e^{\nu(t,r)} dt) = \nu' e^{\nu(t,r)} dr \wedge dt = \nu' e^{-\lambda(t,r)} \omega^{\hat{r}} \wedge \omega^{\hat{t}} = -\nu' e^{-\lambda(t,r)} \omega^{\hat{t}} \wedge \omega^{\hat{r}} \\ &= -\Gamma_{\hat{r}\hat{t}} \wedge \omega^{\hat{r}} \\ d\omega^{\hat{r}} &= d(e^{\lambda(t,r)} dr) = \lambda e^{\lambda(t,r)} dt \wedge dr = \lambda e^{-\nu(t,r)} \omega^{\hat{t}} \wedge \omega^{\hat{r}} = -\lambda e^{-\nu(t,r)} \omega^{\hat{r}} \wedge \omega^{\hat{t}} \\ &= -\Gamma_{\hat{t}\hat{r}} \wedge \omega^{\hat{t}} \\ d\omega^{\hat{\theta}} &= d(r d\theta) = dr \wedge d\theta = \frac{1}{r} e^{-\lambda(r,t)} \omega^{\hat{r}} \wedge \omega^{\hat{\theta}} = -\frac{1}{r} e^{-\lambda(r,t)} \omega^{\hat{\theta}} \wedge \omega^{\hat{r}} = -\Gamma_{\hat{r}\hat{\theta}} \wedge \omega^{\hat{r}} \\ d\omega^{\hat{\phi}} &= d(r \sin \theta d\phi) = \sin \theta dr \wedge d\phi + r \cos \theta d\theta \wedge d\phi \\ &= \frac{1}{r} e^{-\lambda(r,t)} \omega^{\hat{r}} \wedge \omega^{\hat{\phi}} + \frac{\cot \theta}{r} \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} = -\frac{1}{r} e^{-\lambda(r,t)} \omega^{\hat{\phi}} \wedge \omega^{\hat{r}} - \frac{\cot \theta}{r} \omega^{\hat{\phi}} \wedge \omega^{\hat{\theta}} \\ &= -\Gamma_{\hat{r}\hat{\phi}} \wedge \omega^{\hat{r}} - \Gamma_{\hat{\theta}\hat{\phi}} \wedge \omega^{\hat{\theta}}\end{aligned}$$

11.3.1.1 Collecting the results

In this case we have to be particular careful in reading off the curvature one forms. The curvature one-forms are antisymmetric: $\Gamma_{\hat{a}\hat{b}} = -\Gamma_{\hat{b}\hat{a}}$. This means that $\Gamma_{\hat{r}\hat{t}} = \eta^{\hat{t}\hat{t}} \Gamma_{\hat{t}\hat{r}} = -\eta^{\hat{t}\hat{t}} \Gamma_{\hat{r}\hat{t}} = -\eta^{\hat{t}\hat{t}} \eta_{\hat{r}\hat{r}} \Gamma^{\hat{r}}_{\hat{t}} = \Gamma^{\hat{r}}_{\hat{t}}$. But in the former calculation we found that $\Gamma^{\hat{t}}_{\hat{r}} = \nu' e^{-\lambda(t,r)} \omega^{\hat{t}}$ and $\Gamma^{\hat{r}}_{\hat{t}} = \lambda e^{-\nu(t,r)} \omega^{\hat{r}}$, which means that we in order to fulfill the antisymmetric properties need to require that $\Gamma^{\hat{t}}_{\hat{r}} = \Gamma^{\hat{r}}_{\hat{t}} = \lambda e^{-\nu(t,r)} \omega^{\hat{r}} + \nu' e^{-\lambda(t,r)} \omega^{\hat{t}}$, because $\Gamma^{\hat{t}}_{\hat{r}} = \nu' e^{-\lambda(t,r)} \omega^{\hat{t}} +$ (something that makes $\Gamma_{\hat{r}\hat{t}}$ antisymmetric), and $\Gamma^{\hat{r}}_{\hat{t}} = \lambda e^{-\nu(t,r)} \omega^{\hat{r}} +$ (something that makes $\Gamma_{\hat{t}\hat{r}}$ antisymmetric).

$$\begin{aligned}\Gamma^{\hat{t}}_{\hat{r}} &= \nu' e^{-\lambda(t,r)} \omega^{\hat{t}} + \text{Symmetric} = \nu' e^{-\lambda(t,r)} \omega^{\hat{t}} + \lambda e^{-\nu(t,r)} \omega^{\hat{r}} \\ \Gamma^{\hat{r}}_{\hat{t}} &= \lambda e^{-\nu(t,r)} \omega^{\hat{r}} + \text{Symmetric} = \nu' e^{-\lambda(t,r)} \omega^{\hat{t}} + \lambda e^{-\nu(t,r)} \omega^{\hat{r}} \\ \Gamma^{\hat{\theta}}_{\hat{r}} &= \frac{1}{r} e^{-\lambda(r,t)} \omega^{\hat{\theta}} \\ \Gamma^{\hat{\phi}}_{\hat{r}} &= \frac{1}{r} e^{-\lambda(r,t)} \omega^{\hat{\phi}} \\ \Gamma^{\hat{\phi}}_{\hat{\theta}} &= \frac{\cot \theta}{r} \omega^{\hat{\phi}}\end{aligned}$$

Summarizing the curvature one forms in a matrix - Where \hat{a} refers to column and \hat{b} to row:

$$\Gamma^{\hat{a}}_{\hat{b}} = \begin{pmatrix} 0 & \lambda e^{-\nu(t,r)} \omega^{\hat{r}} + \nu' e^{-\lambda(t,r)} \omega^{\hat{t}} & 0 & 0 \\ S & 0 & \frac{1}{r} e^{-\lambda(r,t)} \omega^{\hat{\theta}} & \frac{1}{r} e^{-\lambda(r,t)} \omega^{\hat{\phi}} \\ 0 & AS & 0 & \frac{\cot \theta}{r} \omega^{\hat{\phi}} \\ 0 & AS & AS & 0 \end{pmatrix}$$

11.3.2 The curvature two forms

$$\Omega^{\hat{a}}_{\hat{b}} = d\Gamma^{\hat{a}}_{\hat{b}} + \Gamma^{\hat{a}}_{\hat{c}} \wedge \Gamma^{\hat{c}}_{\hat{b}} = \frac{1}{2} R^{\hat{a}}_{\hat{b}\hat{c}\hat{d}} \omega^{\hat{c}} \wedge \omega^{\hat{d}}$$

We need

$$\begin{aligned} d\Gamma^{\hat{r}}_{\hat{t}} &= d(\lambda e^{-v(t,r)} \omega^{\hat{r}} + v' e^{-\lambda(t,r)} \omega^{\hat{t}}) \\ &= d(\lambda e^{\lambda(t,r)-v(t,r)} dr + v' e^{v(t,r)-\lambda(t,r)} dt) \\ &= (\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{v})) e^{\lambda(t,r)-v(t,r)} dt \wedge dr + (v'' + v'(v' - \lambda')) e^{v(t,r)-\lambda(t,r)} dr \wedge dt \\ &= \left[-(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{v})) e^{-2v(t,r)} + (v'' + v'(v' - \lambda')) e^{-2\lambda(t,r)} \right] \omega^{\hat{r}} \wedge \omega^{\hat{t}} \end{aligned}$$

$$d\Gamma^{\hat{\theta}}_{\hat{t}} = 0$$

$$d\Gamma^{\hat{\phi}}_{\hat{t}} = 0$$

$$\begin{aligned} d\Gamma^{\hat{\theta}}_{\hat{r}} &= d\left(\frac{1}{r} e^{-\lambda(t,r)} \omega^{\hat{\theta}}\right) = d(e^{-\lambda(t,r)} d\theta) = -\dot{\lambda} e^{-\lambda(t,r)} dt \wedge d\theta - \lambda' e^{-\lambda(t,r)} dr \wedge d\theta \\ &= \frac{\dot{\lambda}}{r} e^{-v(t,r)-\lambda(t,r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{t}} + \frac{\lambda'}{r} e^{-2\lambda(t,r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{r}} \end{aligned}$$

$$\begin{aligned} d\Gamma^{\hat{\phi}}_{\hat{r}} &= d\left(\frac{1}{r} e^{-\lambda(t,r)} \omega^{\hat{\phi}}\right) = d(e^{-\lambda(t,r)} \sin \theta d\phi) \\ &= -\dot{\lambda} e^{-\lambda(t,r)} \sin \theta dt \wedge d\phi - \lambda' e^{-\lambda(t,r)} \sin \theta dr \wedge d\phi + e^{-\lambda(t,r)} \cos \theta d\theta \wedge d\phi \\ &= \frac{\dot{\lambda}}{r} e^{-v(t,r)-\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{t}} + \frac{\lambda'}{r} e^{-2\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{r}} + \frac{1}{r^2} e^{-\lambda(t,r)} \cot \theta \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} \end{aligned}$$

$$d\Gamma^{\hat{\phi}}_{\hat{\theta}} = d\left(\frac{\cot \theta}{r} \omega^{\hat{\phi}}\right) = d(\cos \theta d\phi) = -\sin \theta d\theta \wedge d\phi = -\frac{1}{r^2} \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}}$$

11.3.2.1 Collecting the results

Summerized i an matrix where \hat{a} corresponds to column and \hat{b} to row

$$d\Gamma^{\hat{a}}_{\hat{b}} = \left\{ \begin{array}{l} \left[\begin{array}{l} -(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{v})) e^{-2v(t,r)} \\ +(v'' + v'(v' - \lambda')) e^{-2\lambda(t,r)} \end{array} \right] \omega^{\hat{r}} \wedge \omega^{\hat{t}} \\ \left[\begin{array}{l} \frac{\dot{\lambda}}{r} e^{-v(t,r)-\lambda(t,r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{t}} \\ + \frac{\lambda'}{r} e^{-2\lambda(t,r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{r}} \end{array} \right] \\ \left[\begin{array}{l} \frac{\dot{\lambda}}{r} e^{-v(t,r)-\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{t}} \\ + \frac{\lambda'}{r} e^{-2\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{r}} \\ + \frac{1}{r^2} e^{-\lambda(t,r)} \cot \theta \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} \\ - \frac{1}{r^2} \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} \end{array} \right] \end{array} \right\}$$

$$\begin{aligned} \Rightarrow \Omega^{\hat{r}}_{\hat{t}} &= d\Gamma^{\hat{r}}_{\hat{t}} + \Gamma^{\hat{r}}_{\hat{\theta}} \wedge \Gamma^{\hat{\theta}}_{\hat{t}} + \Gamma^{\hat{r}}_{\hat{\phi}} \wedge \Gamma^{\hat{\phi}}_{\hat{t}} \\ &= \left[-(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{v})) e^{-2v(t,r)} + (v'' + v'(v' - \lambda')) e^{-2\lambda(t,r)} \right] \omega^{\hat{r}} \wedge \omega^{\hat{t}} \\ \Omega^{\hat{\theta}}_{\hat{t}} &= d\Gamma^{\hat{\theta}}_{\hat{t}} + \Gamma^{\hat{\theta}}_{\hat{r}} \wedge \Gamma^{\hat{r}}_{\hat{t}} + \Gamma^{\hat{\theta}}_{\hat{\phi}} \wedge \Gamma^{\hat{\phi}}_{\hat{t}} = \Gamma^{\hat{\theta}}_{\hat{r}} \wedge \Gamma^{\hat{r}}_{\hat{t}} \\ &= \frac{1}{r} e^{-\lambda(t,r)} \omega^{\hat{\theta}} \wedge (\lambda e^{-v(t,r)} \omega^{\hat{r}} + v' e^{-\lambda(t,r)} \omega^{\hat{t}}) \\ &= \frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-v(t,r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{r}} + \frac{v'}{r} e^{-2\lambda(t,r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{t}} \\ \Omega^{\hat{\phi}}_{\hat{t}} &= d\Gamma^{\hat{\phi}}_{\hat{t}} + \Gamma^{\hat{\phi}}_{\hat{r}} \wedge \Gamma^{\hat{r}}_{\hat{t}} + \Gamma^{\hat{\phi}}_{\hat{\theta}} \wedge \Gamma^{\hat{\theta}}_{\hat{t}} = \Gamma^{\hat{\phi}}_{\hat{r}} \wedge \Gamma^{\hat{r}}_{\hat{t}} \\ &= \frac{1}{r} e^{-\lambda(t,r)} \omega^{\hat{\phi}} \wedge (\lambda e^{-v(t,r)} \omega^{\hat{r}} + v' e^{-\lambda(t,r)} \omega^{\hat{t}}) \\ &= \frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-v(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{r}} + \frac{v'}{r} e^{-2\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{t}} \end{aligned}$$

$$\begin{aligned}
\Omega^{\hat{\theta}}_{\hat{r}} &= d\Gamma^{\hat{\theta}}_{\hat{r}} + \Gamma^{\hat{\theta}}_{\hat{t}} \wedge \Gamma^{\hat{t}}_{\hat{r}} + \Gamma^{\hat{\theta}}_{\hat{\phi}} \wedge \Gamma^{\hat{\phi}}_{\hat{r}} \\
&= \frac{\dot{\lambda}}{r} e^{-v(t,r)-\lambda(t,r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{t}} + \frac{\lambda'}{r} e^{-2\lambda(t,r)} \omega^{\hat{\theta}} \wedge \omega^{\hat{r}} \\
\Omega^{\hat{\phi}}_{\hat{r}} &= d\Gamma^{\hat{\phi}}_{\hat{r}} + \Gamma^{\hat{\phi}}_{\hat{t}} \wedge \Gamma^{\hat{t}}_{\hat{r}} + \Gamma^{\hat{\phi}}_{\hat{\theta}} \wedge \Gamma^{\hat{\theta}}_{\hat{r}} \\
\Omega^{\hat{\phi}}_{\hat{\theta}} &= d\Gamma^{\hat{\phi}}_{\hat{\theta}} + \Gamma^{\hat{\phi}}_{\hat{t}} \wedge \Gamma^{\hat{t}}_{\hat{\theta}} + \Gamma^{\hat{\phi}}_{\hat{r}} \wedge \Gamma^{\hat{r}}_{\hat{\theta}} \\
&= \frac{\dot{\lambda}}{r} e^{-v(t,r)-\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{t}} + \frac{\lambda'}{r} e^{-2\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{r}} \\
&= -\frac{1}{r^2} \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} - \frac{1}{r^2} e^{-2\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{\theta}} = \frac{(1 - e^{-2\lambda(t,r)})}{r^2} \omega^{\hat{\phi}} \wedge \omega^{\hat{\theta}}
\end{aligned}$$

The diagonal elements:

$$\Omega^{\hat{a}}_{\hat{a}} = d\Gamma^{\hat{a}}_{\hat{a}} + \Gamma^{\hat{a}}_{\hat{c}} \wedge \Gamma^{\hat{c}}_{\hat{a}} = 0$$

11.3.2.2 Collecting the results

Summerized i a matrix where \hat{a} corresponds to column and \hat{b} to row

$$\Omega^{\hat{a}}_{\hat{b}} = \left\{ \begin{array}{c} \left[\begin{array}{c} -(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{v})) e^{-2v(t,r)} \\ +(\nu'' + \nu'(\nu' - \lambda')) e^{-2\lambda(t,r)} \end{array} \right] \omega^{\hat{r}} \wedge \omega^{\hat{t}} \\ AS \\ AS \\ AS \\ \frac{(1 - e^{-2\lambda(t,r)})}{r^2} \omega^{\hat{\phi}} \wedge \omega^{\hat{\theta}} \end{array} \right\}$$

11.3.3 The independent elements of the Riemann tensor in the coordinate basis

$$\begin{aligned}
R^{\hat{r}}_{\hat{t}\hat{r}\hat{t}} &= -(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{v})) e^{-2v(t,r)} + (\nu'' + \nu'(\nu' - \lambda')) e^{-2\lambda(t,r)} \\
R^{\hat{\theta}}_{\hat{t}\hat{\theta}\hat{t}} &= \frac{\nu'}{r} e^{-2\lambda(t,r)} \\
R^{\hat{\theta}}_{\hat{t}\hat{\theta}\hat{r}} &= \frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-v(t,r)} \\
R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{r}} &= \frac{\lambda'}{r} e^{-2\lambda(t,r)} \\
R^{\hat{\phi}}_{\hat{t}\hat{\phi}\hat{t}} &= \frac{\nu'}{r} e^{-2\lambda(t,r)} \\
R^{\hat{\phi}}_{\hat{t}\hat{\phi}\hat{r}} &= \frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-v(t,r)} \\
R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{r}} &= \frac{\lambda'}{r} e^{-2\lambda(t,r)} \\
R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{\theta}} &= \frac{(1 - e^{-2\lambda(t,r)})}{r^2}
\end{aligned}$$

11.3.4 The Ricci tensor:

$$\begin{aligned}
R_{\hat{a}\hat{b}} &= R^{\hat{c}}_{\hat{a}\hat{c}\hat{b}} \\
R_{\hat{t}\hat{t}} &= R^{\hat{c}}_{\hat{t}\hat{c}\hat{t}} = R^{\hat{r}}_{\hat{t}\hat{r}\hat{t}} + R^{\hat{\theta}}_{\hat{t}\hat{\theta}\hat{t}} + R^{\hat{\phi}}_{\hat{t}\hat{\phi}\hat{t}}
\end{aligned}$$

$$7 = \frac{\dot{\lambda}}{r} e^{-v(t,r)-\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{t}} + \frac{\lambda'}{r} e^{-2\lambda(t,r)} \omega^{\hat{\phi}} \wedge \omega^{\hat{r}} + \frac{1}{r^2} e^{-\lambda(t,r)} \cot \theta \omega^{\hat{\theta}} \wedge \omega^{\hat{\phi}} + \frac{\cot \theta}{r} \omega^{\hat{\phi}} \wedge \frac{1}{r} e^{-\lambda(r)} \omega^{\hat{\theta}} =$$

$$\begin{aligned}
 &= -\left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{\nu})\right) e^{-2\nu(t,r)} + \left(\nu'' + \nu'(\nu' - \lambda') + 2\frac{\nu'}{r}\right) e^{-2\lambda(t,r)} \\
 R_{\hat{r}\hat{t}} &= R^{\hat{c}}_{\hat{r}\hat{c}\hat{t}} = R^{\hat{t}}_{\hat{r}\hat{t}\hat{t}} + R^{\hat{r}}_{\hat{r}\hat{r}\hat{t}} + R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{t}} + R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{t}} = R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{t}} + R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{t}} = 2\frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-\nu(t,r)} \\
 &= R_{\hat{t}\hat{r}} \\
 R_{\hat{\theta}\hat{t}} &= R^{\hat{c}}_{\hat{\theta}\hat{c}\hat{t}} = R^{\hat{t}}_{\hat{\theta}\hat{t}\hat{t}} + R^{\hat{r}}_{\hat{\theta}\hat{r}\hat{t}} + R^{\hat{\theta}}_{\hat{\theta}\hat{\theta}\hat{t}} + R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{t}} = 0 \\
 R_{\hat{\phi}\hat{t}} &= R^{\hat{c}}_{\hat{\phi}\hat{c}\hat{t}} = R^{\hat{t}}_{\hat{\phi}\hat{t}\hat{t}} + R^{\hat{r}}_{\hat{\phi}\hat{r}\hat{t}} + R^{\hat{\theta}}_{\hat{\phi}\hat{\theta}\hat{t}} + R^{\hat{\phi}}_{\hat{\phi}\hat{\phi}\hat{t}} = 0 \\
 R_{\hat{r}\hat{r}} &= R^{\hat{c}}_{\hat{r}\hat{c}\hat{r}} = R^{\hat{t}}_{\hat{r}\hat{t}\hat{r}} + R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{r}} + R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{r}} = -R^{\hat{r}}_{\hat{t}\hat{r}\hat{t}} + R^{\hat{\theta}}_{\hat{r}\hat{\theta}\hat{r}} + R^{\hat{\phi}}_{\hat{r}\hat{\phi}\hat{r}} \\
 &= \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{\nu})\right) e^{-2\nu(t,r)} - \left(\nu'' + \nu'(\nu' - \lambda') - 2\frac{\lambda'}{r}\right) e^{-2\lambda(t,r)} \\
 R_{\hat{\theta}\hat{r}} &= R^{\hat{c}}_{\hat{\theta}\hat{c}\hat{r}} = R^{\hat{t}}_{\hat{\theta}\hat{t}\hat{r}} + R^{\hat{r}}_{\hat{\theta}\hat{r}\hat{r}} + R^{\hat{\theta}}_{\hat{\theta}\hat{\theta}\hat{r}} + R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{r}} = 0 \\
 R_{\hat{\phi}\hat{r}} &= R^{\hat{c}}_{\hat{\phi}\hat{c}\hat{r}} = 0 \\
 R_{\hat{\theta}\hat{\theta}} &= R^{\hat{c}}_{\hat{\theta}\hat{c}\hat{\theta}} = R^{\hat{t}}_{\hat{\theta}\hat{t}\hat{\theta}} + R^{\hat{r}}_{\hat{\theta}\hat{r}\hat{\theta}} + R^{\hat{\theta}}_{\hat{\theta}\hat{\theta}\hat{\theta}} + R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{\theta}} = -R^{\hat{\theta}}_{\hat{t}\hat{\theta}\hat{t}} + R^{\hat{r}}_{\hat{r}\hat{\theta}\hat{r}} + R^{\hat{\phi}}_{\hat{\theta}\hat{\phi}\hat{\theta}} \\
 &= \left(-\frac{\nu'}{r} + \frac{\lambda'}{r}\right) e^{-2\lambda(t,r)} + \frac{(1 - e^{-2\lambda(t,r)})}{r^2} \\
 R_{\hat{\phi}\hat{\theta}} &= R^{\hat{c}}_{\hat{\phi}\hat{c}\hat{\theta}} = 0 \\
 R_{\hat{\phi}\hat{\phi}} &= R^{\hat{c}}_{\hat{\phi}\hat{c}\hat{\phi}} = R^{\hat{t}}_{\hat{\phi}\hat{t}\hat{\phi}} + R^{\hat{r}}_{\hat{\phi}\hat{r}\hat{\phi}} + R^{\hat{\theta}}_{\hat{\phi}\hat{\theta}\hat{\phi}} + R^{\hat{\phi}}_{\hat{\phi}\hat{\phi}\hat{\phi}} = -R^{\hat{\phi}}_{\hat{t}\hat{\phi}\hat{t}} + R^{\hat{r}}_{\hat{r}\hat{\phi}\hat{r}} + R^{\hat{\theta}}_{\hat{\theta}\hat{\phi}\hat{\theta}} \\
 &= \left(-\frac{\nu'}{r} + \frac{\lambda'}{r}\right) e^{-2\lambda(t,r)} + \frac{(1 - e^{-2\lambda(t,r)})}{r^2}
 \end{aligned}$$

Collecting the results: Summerized i an matrix where \hat{a} corresponds to column and \hat{b} to row

$$R_{\hat{a}\hat{b}} = \begin{pmatrix} \begin{bmatrix} -(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{\nu})) e^{-2\nu(t,r)} \\ +(\nu'' + \nu'(\nu' - \lambda') + 2\frac{\nu'}{r}) e^{-2\lambda(t,r)} \end{bmatrix} & 2\frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-\nu(t,r)} \\ S & \begin{bmatrix} (\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{\nu})) e^{-2\nu(t,r)} \\ -(\nu'' + \nu'(\nu' - \lambda') - 2\frac{\lambda'}{r}) e^{-2\lambda(t,r)} \end{bmatrix} \\ & \begin{bmatrix} \left(-\frac{\nu'}{r} + \frac{\lambda'}{r}\right) e^{-2\lambda(t,r)} \\ +\frac{(1 - e^{-2\lambda(t,r)})}{r^2} \end{bmatrix} \\ & \begin{bmatrix} \left(-\frac{\nu'}{r} + \frac{\lambda'}{r}\right) e^{-2\lambda(t,r)} \\ +\frac{(1 - e^{-2\lambda(t,r)})}{r^2} \end{bmatrix} \end{pmatrix}$$

Notice:

$$\begin{aligned}
 R_{\hat{t}\hat{t}} + R_{\hat{r}\hat{r}} &= 2\left(\frac{\nu'}{r} + \frac{\lambda'}{r}\right) e^{-2\lambda(t,r)} \\
 R_{\hat{\theta}\hat{\theta}} &= R_{\hat{\phi}\hat{\phi}}
 \end{aligned}$$

11.3.5 The Ricci scalar

$$R = \eta^{\hat{a}\hat{b}} R_{\hat{a}\hat{b}} = \eta^{\hat{t}\hat{t}} R_{\hat{t}\hat{t}} + \eta^{\hat{r}\hat{r}} R_{\hat{r}\hat{r}} + \eta^{\hat{\theta}\hat{\theta}} R_{\hat{\theta}\hat{\theta}} + \eta^{\hat{\phi}\hat{\phi}} R_{\hat{\phi}\hat{\phi}} = R_{\hat{t}\hat{t}} - R_{\hat{r}\hat{r}} - 2R_{\hat{\theta}\hat{\theta}}$$

$$= {}^8 - 2 \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v}) \right) e^{-2v(t,r)} + 2 \left(v'' + v'(v' - \lambda') + \frac{1}{r^2} + 2 \left(\frac{v'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} - 2 \frac{1}{r^2}$$

11.3.6 The Einstein tensor

$$\begin{aligned} G_{\hat{a}\hat{b}} &= R_{\hat{a}\hat{b}} - \frac{1}{2} \eta_{\hat{a}\hat{b}} R \\ \Rightarrow G_{\hat{t}\hat{t}} &= R_{\hat{t}\hat{t}} - \frac{1}{2} \eta_{\hat{t}\hat{t}} R = {}^9 \left(2 \frac{\lambda'}{r} - \frac{1}{r^2} \right) e^{-2\lambda(t,r)} + \frac{1}{r^2} \\ G_{\hat{r}\hat{t}} &= R_{\hat{r}\hat{t}} - \frac{1}{2} \eta_{\hat{r}\hat{t}} R = R_{\hat{r}\hat{t}} = 2 \frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-v(t,r)} = G_{\hat{t}\hat{r}} \\ G_{\hat{r}\hat{r}} &= R_{\hat{r}\hat{r}} - \frac{1}{2} \eta_{\hat{r}\hat{r}} R = R_{\hat{r}\hat{r}} + \frac{1}{2} R = {}^{10} \left(\frac{1}{r^2} + 2 \frac{v'}{r} \right) e^{-2\lambda(t,r)} - \frac{1}{r^2} \\ G_{\hat{\theta}\hat{\theta}} &= R_{\hat{\theta}\hat{\theta}} - \frac{1}{2} \eta_{\hat{\theta}\hat{\theta}} R \\ &= R_{\hat{\theta}\hat{\theta}} + \frac{1}{2} R \\ &= {}^{11} - \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v}) \right) e^{-2v(t,r)} + \left(v'' + v'(v' - \lambda') + \left(\frac{v'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} \\ &= G_{\hat{\phi}\hat{\phi}} \end{aligned}$$

Collecting the results: Summnerized i an matrix where \hat{a} corresponds to column and \hat{b} to row

$$\begin{aligned} {}^8 &= - \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v}) \right) e^{-2v(t,r)} + \left(v'' + v'(v' - \lambda') + 2 \frac{v'}{r} \right) e^{-2\lambda(t,r)} - \left(\left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v}) \right) e^{-2v(t,r)} - \right. \\ &\quad \left. \left(v'' + v'(v' - \lambda') - 2 \frac{\lambda'}{r} \right) e^{-2\lambda(t,r)} \right) - 2 \left(\left(-\frac{v'}{r} + \frac{\lambda'}{r} \right) e^{-2\lambda(t,r)} + \frac{(1-e^{-2\lambda(t,r)})}{r^2} \right) = \\ {}^9 &= - \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v}) \right) e^{-2v(t,r)} + \left(v'' + v'(v' - \lambda') + 2 \frac{v'}{r} \right) e^{-2\lambda(t,r)} - \frac{1}{2} \left(-2 \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v}) \right) e^{-2v(t,r)} + \right. \\ &\quad \left. 2 \left(v'' + v'(v' - \lambda') + \frac{1}{r^2} + 2 \left(\frac{v'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} - 2 \frac{1}{r^2} \right) = \left(v'' + v'(v' - \lambda') + 2 \frac{v'}{r} \right) e^{-2\lambda(t,r)} - \\ &\quad \left(v'' + v'(v' - \lambda') + \frac{1}{r^2} + 2 \left(\frac{v'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} + \frac{1}{r^2} = \\ {}^{10} &= \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v}) \right) e^{-2v(t,r)} - \left(v'' + v'(v' - \lambda') - 2 \frac{\lambda'}{r} \right) e^{-2\lambda(t,r)} + \frac{1}{2} \left(-2 \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v}) \right) e^{-2v(t,r)} + \right. \\ &\quad \left. 2 \left(v'' + v'(v' - \lambda') + \frac{1}{r^2} + 2 \left(\frac{v'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} - 2 \frac{1}{r^2} \right) = - \left(v'' + v'(v' - \lambda') - 2 \frac{\lambda'}{r} \right) e^{-2\lambda(t,r)} + \\ &\quad \left(v'' + v'(v' - \lambda') + \frac{1}{r^2} + 2 \left(\frac{v'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} - \frac{1}{r^2} = \\ {}^{11} &= \left(-\frac{v'}{r} + \frac{\lambda'}{r} \right) e^{-2\lambda(t,r)} + \frac{(1-e^{-2\lambda(t,r)})}{r^2} + \frac{1}{2} \left(-2 \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v}) \right) e^{-2v(t,r)} + 2 \left(v'' + v'(v' - \lambda') + \frac{1}{r^2} + \right. \right. \\ &\quad \left. \left. 2 \left(\frac{v'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} - 2 \frac{1}{r^2} \right) = \end{aligned}$$

$$G_{\hat{a}\hat{b}} = \begin{pmatrix} \left[\begin{array}{cc} \left(2\frac{\lambda'}{r} - \frac{1}{r^2} \right) e^{-2\lambda(t,r)} & 2\frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-\nu(t,r)} \\ +\frac{1}{r^2} & \end{array} \right] & \\ S & \left[\begin{array}{cc} \left(\frac{1}{r^2} + 2\frac{\nu'}{r} \right) e^{-2\lambda(t,r)} & \\ & -\frac{1}{r^2} \end{array} \right] \\ & \left[\begin{array}{c} -(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)} \\ + \left(\nu'' + \nu'(\nu' - \lambda') + \left(\frac{\nu'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} \end{array} \right] \\ & \left[\begin{array}{c} -(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)} \\ + \left(\nu'' + \nu'(\nu' - \lambda') + \left(\frac{\nu'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} \end{array} \right] \end{pmatrix}$$

11.3.7 The Ricci and Einstein tensor in the coordinate basis:

The Ricci tensor

$$\begin{aligned} R_{ab} &= \Lambda^{\hat{c}}_a \Lambda^{\hat{d}}_b R_{\hat{c}\hat{d}} \\ \Lambda^{\hat{c}}_a &= \begin{pmatrix} e^{\nu(t,r)} & & & \\ & e^{\lambda(t,r)} & & \\ & & r & \\ & & & r \sin \theta \end{pmatrix} \\ \Rightarrow R_{tt} &= \Lambda^{\hat{c}}_t \Lambda^{\hat{d}}_t R_{\hat{c}\hat{d}} = (\Lambda^{\hat{t}}_t)^2 R_{\hat{t}\hat{t}} \\ &= e^{2\nu(t,r)} \left(-(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)} + \left(\nu'' + \nu'(\nu' - \lambda') + 2\frac{\nu'}{r} \right) e^{-2\lambda(t,r)} \right) \\ &= -(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) + \left(\nu'' + \nu'(\nu' - \lambda') + 2\frac{\nu'}{r} \right) e^{2\nu(t,r)-2\lambda(t,r)} \\ R_{rt} &= \Lambda^{\hat{c}}_r \Lambda^{\hat{d}}_t R_{\hat{c}\hat{d}} = \Lambda^{\hat{r}}_r \Lambda^{\hat{t}}_t R_{\hat{r}\hat{t}} = e^{\lambda(t,r)} e^{\nu(t,r)} 2\frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-\nu(t,r)} = 2\frac{\dot{\lambda}}{r} \\ R_{rr} &= \Lambda^{\hat{c}}_r \Lambda^{\hat{d}}_r R_{\hat{c}\hat{d}} = (\Lambda^{\hat{r}}_r)^2 R_{\hat{r}\hat{r}} \\ &= e^{2\lambda(t,r)} \left((\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)} - \left(\nu'' + \nu'(\nu' - \lambda') - 2\frac{\lambda'}{r} \right) e^{-2\lambda(t,r)} \right) \\ &= (\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)+2\lambda(t,r)} - \left(\nu'' + \nu'(\nu' - \lambda') - 2\frac{\lambda'}{r} \right) \\ R_{\theta\theta} &= \Lambda^{\hat{c}}_\theta \Lambda^{\hat{d}}_\theta R_{\hat{c}\hat{d}} = (\Lambda^{\hat{\theta}}_\theta)^2 R_{\hat{\theta}\hat{\theta}} \\ &= r^2 \left(\left(-\frac{\nu'}{r} + \frac{\lambda'}{r} \right) e^{-2\lambda(t,r)} + \frac{(1 - e^{-2\lambda(t,r)})}{r^2} \right) \\ &= ((-\nu' + \lambda')r - 1) e^{-2\lambda(t,r)} + 1 \\ R_{\phi\phi} &= \Lambda^{\hat{c}}_\phi \Lambda^{\hat{d}}_\phi R_{\hat{c}\hat{d}} = (\Lambda^{\hat{\phi}}_\phi)^2 R_{\hat{\phi}\hat{\phi}} \\ &= r^2 \sin^2 \theta \left(\left(-\frac{\nu'}{r} + \frac{\lambda'}{r} \right) e^{-2\lambda(t,r)} + \frac{(1 - e^{-2\lambda(t,r)})}{r^2} \right) \\ &= (((-\nu' + \lambda')r - 1) e^{-2\lambda(t,r)} + 1) \sin^2 \theta \end{aligned}$$

Collecting the results:

$$R_{tt} = -(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) + \left(\nu'' + \nu'(\nu' - \lambda') + 2\frac{\nu'}{r} \right) e^{2\nu(t,r)-2\lambda(t,r)}$$

$$\begin{aligned}
R_{rt} &= R_{tr} = 2\frac{\dot{\lambda}}{r} \\
R_{rr} &= \left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v})\right)e^{-2\nu(t,r)+2\lambda(t,r)} - \left(v'' + v'(v' - \lambda') - 2\frac{\lambda'}{r}\right) \\
R_{\theta\theta} &= \left((-v' + \lambda')r - 1\right)e^{-2\lambda(t,r)} + 1 \\
R_{\phi\phi} &= \left(\left((-v' + \lambda')r - 1\right)e^{-2\lambda(t,r)} + 1\right)\sin^2\theta = R_{\theta\theta}\sin^2\theta
\end{aligned}$$

11.3.8 The Einstein tensor

$$\begin{aligned}
G_{ab} &= \Lambda^c_a \Lambda^{\hat{d}}_b G_{\hat{c}\hat{d}} \\
\Lambda^{\hat{c}}_a &= \left\{ \begin{array}{ccc} e^{\nu(t,r)} & & \\ & e^{\lambda(t,r)} & \\ & & r \\ & & & r \sin\theta \end{array} \right\} \\
\Rightarrow G_{tt} &= \Lambda^{\hat{c}}_t \Lambda^{\hat{d}}_t G_{\hat{c}\hat{d}} = (\Lambda^{\hat{t}}_t)^2 G_{\hat{t}\hat{t}} = e^{2\nu(t,r)} \left(\left(2\frac{\lambda'}{r} - \frac{1}{r^2} \right) e^{-2\lambda(t,r)} + \frac{1}{r^2} \right) \\
G_{rt} &= \Lambda^{\hat{c}}_r \Lambda^{\hat{d}}_t G_{\hat{c}\hat{d}} = \Lambda^{\hat{r}}_r \Lambda^{\hat{t}}_t G_{\hat{r}\hat{t}} = e^{\lambda(t,r)} e^{\nu(t,r)} 2\frac{\dot{\lambda}}{r} e^{-\lambda(t,r)-\nu(t,r)} = 2\frac{\dot{\lambda}}{r} \\
G_{rr} &= \Lambda^{\hat{c}}_r \Lambda^{\hat{d}}_r G_{\hat{c}\hat{d}} = (\Lambda^{\hat{r}}_r)^2 G_{\hat{r}\hat{r}} \\
&= e^{2\lambda(t,r)} \left(\left(\frac{1}{r^2} + 2\frac{v'}{r} \right) e^{-2\lambda(t,r)} - \frac{1}{r^2} \right) = \left(\frac{1}{r^2} + 2\frac{v'}{r} \right) - \frac{e^{2\lambda(t,r)}}{r^2} \\
G_{\theta\theta} &= \Lambda^{\hat{c}}_\theta \Lambda^{\hat{d}}_\theta G_{\hat{c}\hat{d}} = (\Lambda^{\hat{\theta}}_\theta)^2 G_{\hat{\theta}\hat{\theta}} \\
&= r^2 \left[-\left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v})\right)e^{-2\nu(t,r)} + \left(v'' + v'(v' - \lambda') + \left(\frac{v'}{r} - \frac{\lambda'}{r}\right)\right)e^{-2\lambda(t,r)} \right] \\
G_{\phi\phi} &= \Lambda^{\hat{c}}_\phi \Lambda^{\hat{d}}_\phi G_{\hat{c}\hat{d}} = (\Lambda^{\hat{\phi}}_\phi)^2 G_{\hat{\phi}\hat{\phi}} \\
&= r^2 \sin^2\theta \left[-\left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v})\right)e^{-2\nu(t,r)} + \left(v'' + v'(v' - \lambda') + \left(\frac{v'}{r} - \frac{\lambda'}{r}\right)\right)e^{-2\lambda(t,r)} \right]
\end{aligned}$$

Collecting the results:

$$\begin{aligned}
G_{tt} &= e^{2\nu(t,r)} \left(\left(2\frac{\lambda'}{r} - \frac{1}{r^2} \right) e^{-2\lambda(t,r)} + \frac{1}{r^2} \right) \\
G_{rt} &= G_{tr} = 2\frac{\dot{\lambda}}{r} \\
G_{rr} &= \frac{1}{r^2} + 2\frac{v'}{r} - \frac{e^{2\lambda(t,r)}}{r^2} \\
G_{\theta\theta} &= r^2 \left[-\left(\ddot{\lambda} + \dot{\lambda}(\dot{\lambda} - \dot{v})\right)e^{-2\nu(t,r)} + \left(v'' + v'(v' - \lambda') + \left(\frac{v'}{r} - \frac{\lambda'}{r}\right)\right)e^{-2\lambda(t,r)} \right] \\
G_{\phi\phi} &= \sin^2\theta G_{\theta\theta}
\end{aligned}$$

11.3.9 The Einstein tensor of second kind

$$G^a_b = g^{ac} G_{bc}$$

$$g^{ab} = \left\{ \begin{array}{ccc} e^{-2\nu(t,r)} & & \\ & -e^{-2\lambda(t,r)} & \\ & & -\frac{1}{r^2} \\ & & & \frac{1}{r^2 \sin^2 \theta} \end{array} \right\}$$

$$\Rightarrow G_t^t = g^{tc} G_{tc} = g^{tt} G_{tt} = e^{-2\nu(t,r)} e^{2\nu(t,r)} \left(\left(2 \frac{\lambda'}{r} - \frac{1}{r^2} \right) e^{-2\lambda(t,r)} + \frac{1}{r^2} \right)$$

$$= \left(2 \frac{\lambda'}{r} - \frac{1}{r^2} \right) e^{-2\lambda(t,r)} + \frac{1}{r^2}$$

$$G_r^t = g^{tc} G_{rc} = g^{tt} G_{rt} = e^{-2\nu(t,r)} 2 \frac{\dot{\lambda}}{r}$$

$$G_t^r = g^{rc} G_{tc} = g^{rr} G_{tr} = -e^{-2\lambda(t,r)} 2 \frac{\dot{\lambda}}{r}$$

$$G_r^r = g^{rr} G_{rr} = -e^{-2\lambda(t,r)} \left(\left(\frac{1}{r^2} + 2 \frac{\nu'}{r} \right) - \frac{e^{2\lambda(t,r)}}{r^2} \right) = \frac{1}{r^2} - e^{-2\lambda(t,r)} \left(\frac{1}{r^2} + 2 \frac{\nu'}{r} \right)$$

$$G_\theta^\theta = g^{\theta\theta} G_{\theta\theta}$$

$$= -\frac{1}{r^2} r^2 \left[-(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)} + \left(\nu'' + \nu'(\nu' - \lambda') + \left(\frac{\nu'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} \right]$$

$$= (\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)} - \left(\nu'' + \nu'(\nu' - \lambda') + \left(\frac{\nu'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)}$$

$$G_\phi^\phi = g^{\phi\phi} G_{\phi\phi}$$

$$= {}^{12} (\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)} - \left(\nu'' + \nu'(\nu' - \lambda') + \left(\frac{\nu'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)}$$

Collecting the results

$$G_t^t = \left(2 \frac{\lambda'}{r} - \frac{1}{r^2} \right) e^{-2\lambda(t,r)} + \frac{1}{r^2}$$

$$G_r^t = e^{-2\nu(t,r)} 2 \frac{\dot{\lambda}}{r}$$

$$G_t^r = -e^{-2\lambda(t,r)} 2 \frac{\dot{\lambda}}{r}$$

$$G_r^r = \frac{1}{r^2} - e^{-2\lambda(t,r)} \left(\frac{1}{r^2} + 2 \frac{\nu'}{r} \right)$$

$$G_\theta^\theta = (\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)} - \left(\nu'' + \nu'(\nu' - \lambda') + \left(\frac{\nu'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)}$$

$$G_\phi^\phi = G_\theta^\theta$$

$${}^{12} = -\frac{1}{r^2 \sin^2 \theta} r^2 \sin^2 \theta \left[-(\ddot{\lambda} + \dot{\lambda}(\lambda - \dot{\nu})) e^{-2\nu(t,r)} + \left(\nu'' + \nu'(\nu' - \lambda') + \left(\frac{\nu'}{r} - \frac{\lambda'}{r} \right) \right) e^{-2\lambda(t,r)} \right] =$$

11.4 Other Examples

11.4.1 Calculation of the scalar $R_{abcd}R^{abcd}$ in the Schwarzschild metric

$$\begin{aligned}
 R_{abcd}R^{abcd} &= R_{\hat{r}\hat{t}\hat{r}\hat{t}}R^{\hat{t}\hat{r}\hat{t}\hat{r}} + R_{\hat{r}\hat{t}\hat{t}\hat{r}}R^{\hat{r}\hat{t}\hat{t}\hat{r}} + R_{\hat{t}\hat{r}\hat{t}\hat{r}}R^{\hat{r}\hat{t}\hat{r}\hat{t}} + R_{\hat{t}\hat{r}\hat{r}\hat{t}}R^{\hat{t}\hat{r}\hat{r}\hat{t}} + R_{\hat{\theta}\hat{t}\hat{\theta}\hat{t}}R^{\hat{t}\hat{\theta}\hat{t}\hat{\theta}} + R_{\hat{\theta}\hat{t}\hat{t}\hat{\theta}}R^{\hat{t}\hat{\theta}\hat{t}\hat{\theta}} \\
 &\quad + R_{\hat{t}\hat{\theta}\hat{t}\hat{\theta}}R^{\hat{t}\hat{\theta}\hat{t}\hat{\theta}} + R_{\hat{t}\hat{\theta}\hat{\theta}\hat{t}}R^{\hat{t}\hat{\theta}\hat{\theta}\hat{t}} + R_{\hat{\theta}\hat{r}\hat{\theta}\hat{r}}R^{\hat{r}\hat{\theta}\hat{r}\hat{\theta}} + R_{\hat{\theta}\hat{r}\hat{r}\hat{\theta}}R^{\hat{r}\hat{\theta}\hat{r}\hat{\theta}} \\
 &\quad + R_{\hat{r}\hat{\theta}\hat{r}\hat{\theta}}R^{\hat{r}\hat{\theta}\hat{r}\hat{\theta}} + R_{\hat{r}\hat{\theta}\hat{\theta}\hat{r}}R^{\hat{r}\hat{\theta}\hat{\theta}\hat{r}} + R_{\hat{\phi}\hat{t}\hat{\phi}\hat{t}}R^{\hat{t}\hat{\phi}\hat{t}\hat{\phi}} + R_{\hat{\phi}\hat{t}\hat{t}\hat{\phi}}R^{\hat{t}\hat{\phi}\hat{t}\hat{\phi}} \\
 &\quad + R_{\hat{t}\hat{\phi}\hat{t}\hat{\phi}}R^{\hat{t}\hat{\phi}\hat{t}\hat{\phi}} + R_{\hat{t}\hat{\phi}\hat{\phi}\hat{t}}R^{\hat{t}\hat{\phi}\hat{\phi}\hat{t}} + R_{\hat{\phi}\hat{r}\hat{\phi}\hat{r}}R^{\hat{r}\hat{\phi}\hat{r}\hat{\phi}} + R_{\hat{\phi}\hat{r}\hat{r}\hat{\phi}}R^{\hat{r}\hat{\phi}\hat{r}\hat{\phi}} \\
 &\quad + R_{\hat{r}\hat{\phi}\hat{r}\hat{\phi}}R^{\hat{r}\hat{\phi}\hat{r}\hat{\phi}} + R_{\hat{r}\hat{\phi}\hat{\phi}\hat{r}}R^{\hat{r}\hat{\phi}\hat{\phi}\hat{r}} + R_{\hat{\phi}\hat{\theta}\hat{\phi}\hat{\theta}}R^{\hat{\theta}\hat{\phi}\hat{\theta}\hat{\phi}} + R_{\hat{\phi}\hat{\theta}\hat{\theta}\hat{\phi}}R^{\hat{\theta}\hat{\phi}\hat{\theta}\hat{\phi}} \\
 &\quad + R_{\hat{\theta}\hat{\phi}\hat{\theta}\hat{\phi}}R^{\hat{\theta}\hat{\phi}\hat{\theta}\hat{\phi}} + R_{\hat{\theta}\hat{\phi}\hat{\phi}\hat{\theta}}R^{\hat{\theta}\hat{\phi}\hat{\phi}\hat{\theta}} \\
 &= \left(\frac{2m}{r^3}\right)\left(\frac{2m}{r^3}\right) + \left(-\frac{2m}{r^3}\right)\left(-\frac{2m}{r^3}\right) + \left(\frac{2m}{r^3}\right)\left(\frac{2m}{r^3}\right) + \left(-\frac{2m}{r^3}\right)\left(-\frac{2m}{r^3}\right) \\
 &\quad + \left(-\frac{m}{r^3}\right)\left(-\frac{m}{r^3}\right) + \left(\frac{m}{r^3}\right)\left(\frac{m}{r^3}\right) + \left(\frac{m}{r^3}\right)\left(\frac{m}{r^3}\right) + \left(-\frac{m}{r^3}\right)\left(-\frac{m}{r^3}\right) \\
 &\quad + \left(\frac{m}{r^3}\right)\left(\frac{m}{r^3}\right) + \left(-\frac{m}{r^3}\right)\left(-\frac{m}{r^3}\right) + \left(\frac{m}{r^3}\right)\left(\frac{m}{r^3}\right) + \left(-\frac{m}{r^3}\right)\left(-\frac{m}{r^3}\right) \\
 &\quad + \left(-\frac{m}{r^3}\right)\left(-\frac{m}{r^3}\right) + \left(\frac{m}{r^3}\right)\left(\frac{m}{r^3}\right) + \left(-\frac{m}{r^3}\right)\left(-\frac{m}{r^3}\right) + \left(\frac{m}{r^3}\right)\left(\frac{m}{r^3}\right) \\
 &\quad + \left(\frac{m}{r^3}\right)\left(\frac{m}{r^3}\right) + \left(-\frac{m}{r^3}\right)\left(-\frac{m}{r^3}\right) + \left(\frac{m}{r^3}\right)\left(\frac{m}{r^3}\right) + \left(-\frac{m}{r^3}\right)\left(-\frac{m}{r^3}\right) \\
 &\quad + \left(-\frac{2m}{r^3}\right)\left(-\frac{2m}{r^3}\right) + \left(\frac{2m}{r^3}\right)\left(\frac{2m}{r^3}\right) + \left(-\frac{2m}{r^3}\right)\left(-\frac{2m}{r^3}\right) + \left(\frac{2m}{r^3}\right)\left(\frac{2m}{r^3}\right) \\
 &= \frac{48m^2}{r^6}
 \end{aligned}$$

11.4.2 Geodesics and Christoffel symbols of the Schwarzschild metric with $\theta = \frac{\pi}{2}$

The line element

$$ds^2 = -\left(1 - \frac{2m}{r}\right)dt^2 + \left(1 - \frac{2m}{r}\right)^{-1}dr^2 + r^2d\phi^2$$

To find the geodesic we use the Euler-Lagrange equation:

$$0 = \frac{d}{ds}\left(\frac{\partial F}{\partial \dot{x}^a}\right) - \frac{\partial F}{\partial x^a}$$

where

$$F = -\left(1 - \frac{2m}{r}\right)\dot{t}^2 + \left(1 - \frac{2m}{r}\right)^{-1}\dot{r}^2 + r^2\dot{\phi}^2$$

$x^a = t$:

$$\frac{\partial F}{\partial t} = 0$$

$$\frac{\partial F}{\partial \dot{t}} = -2\left(1 - \frac{2m}{r}\right)\dot{t}$$

$$\frac{d}{ds}\left(\frac{\partial F}{\partial \dot{t}}\right) = -\frac{4m}{r^2}\dot{r}\dot{t} - 2\left(1 - \frac{2m}{r}\right)\ddot{t}$$

$$\Rightarrow 0 = \frac{4m}{r^2}\dot{r}\dot{t} + 2\left(1 - \frac{2m}{r}\right)\ddot{t}$$

$$\Leftrightarrow 0 = \ddot{t} + \frac{2m}{r(r-2m)}\dot{r}\dot{t}$$

$x^a = r$:

$$\begin{aligned} \frac{\partial F}{\partial r} &= {}^{13} - \frac{2m}{r^2} \dot{t}^2 - \frac{2m}{r^2} \left(1 - \frac{2m}{r}\right)^{-2} \dot{r}^2 + 2r\dot{\phi}^2 \\ \frac{\partial F}{\partial \dot{r}} &= 2 \left(1 - \frac{2m}{r}\right)^{-1} \dot{r} \\ \frac{d}{ds} \left(\frac{\partial F}{\partial \dot{r}} \right) &= 2 \left(1 - \frac{2m}{r}\right)^{-1} \ddot{r} - \frac{4m}{r^2} \left(1 - \frac{2m}{r}\right)^{-2} \dot{r}^2 \\ \Rightarrow 0 &= 2 \left(1 - \frac{2m}{r}\right)^{-1} \ddot{r} - \frac{4m}{r^2} \left(1 - \frac{2m}{r}\right)^{-2} \dot{r}^2 + \frac{2m}{r^2} \dot{t}^2 - 2r\dot{\phi}^2 \\ \Leftrightarrow 0 &= \ddot{r} - \frac{m}{r(r-2m)} \dot{r}^2 + (r-2m) \frac{m}{r^3} \dot{t}^2 - (r-2m)\dot{\phi}^2 \end{aligned}$$

$x^a = \phi$:

$$\begin{aligned} \frac{\partial F}{\partial \phi} &= 0 \\ \frac{\partial F}{\partial \dot{\phi}} &= 2r^2 \dot{\phi} \\ \frac{d}{ds} \left(\frac{\partial F}{\partial \dot{\phi}} \right) &= 4r\dot{r}\dot{\phi} + 2r^2 \ddot{\phi} \\ \Rightarrow 0 &= 4r\dot{r}\dot{\phi} + 2r^2 \ddot{\phi} \\ \Leftrightarrow 0 &= \ddot{\phi} + \frac{2}{r} \dot{r}\dot{\phi} \end{aligned}$$

Collecting the results

$$\begin{aligned} 0 &= \ddot{t} + \frac{2m}{r(r-2m)} \dot{r}\dot{t} \\ 0 &= \ddot{r} - \frac{m}{r(r-2m)} \dot{r}^2 + (r-2m) \frac{m}{r^3} \dot{t}^2 - (r-2m)\dot{\phi}^2 \\ 0 &= \ddot{\phi} + \frac{2}{r} \dot{r}\dot{\phi} \end{aligned}$$

We can now find the Christoffel symbols:

$$\begin{aligned} \Gamma_{rt}^t &= \frac{m}{r(r-2m)} \\ \Gamma_{rr}^r &= -\frac{m}{r(r-2m)} \\ \Gamma_{tt}^r &= \frac{m(r-2m)}{r^3} \\ \Gamma_{\phi\phi}^r &= -(r-2m) \\ \Gamma_{r\phi}^\phi &= \Gamma_{\phi r}^\phi = \frac{1}{r} \end{aligned}$$

11.4.3 The general Schwarzschild metric with nonzero cosmological constant.

11.4.3.1 *The Ricci rotation coefficients and Ricci tensor for the Schwarzschild metric with nonzero cosmological constant.*

The line element

$$ds^2 = f(r)dt^2 - \frac{1}{f(r)}dr^2 - r^2d\theta^2 - r^2\sin^2\theta d\phi^2$$

¹³ $\frac{\partial}{\partial t} \left(\left(1 - \frac{2m}{r}\right)^{-1} \right) = -\frac{2m}{r^2} \left(1 - \frac{2m}{r}\right)^{-2} = -\frac{2m}{(r-2m)^2}$

$$f(r) = 1 - \frac{2m}{r} - \frac{1}{3}\Lambda r^2$$

Now we can compare with the line element of the Schwarzschild metric with zero cosmological constant, where the primes should not be mistaken for the derivative d/dr .

$$ds^2 = e^{2\nu(r')} dt'^2 - e^{2\lambda(r')} dr'^2 - r'^2 d\theta'^2 - r'^2 \sin^2 \theta' d\phi'^2$$

And choose:

$$\begin{aligned} e^{\nu(r')} dt' &= \sqrt{f(r)} dt \\ e^{\lambda(r')} dr' &= \frac{1}{\sqrt{f(r)}} dr \\ r' d\theta' &= r d\theta \\ r' \sin \theta' d\phi' &= r \sin \theta d\phi \end{aligned}$$

Comparing the two metrics we see: $\phi' = \phi, \theta' = \theta, r' = r, e^{\nu(r')} = \sqrt{f(r)}, \nu = -\lambda, t' = t$

Next we can use the former calculations of the Schwarzschild metric with zero cosmological constant to find the Ricci rotation coefficients and the Ricci tensor for the Schwarzschild metric with non-zero cosmological constant.

$$\begin{aligned} \Gamma^{\hat{t}}_{\hat{r}\hat{t}} &= \Gamma^{\hat{r}}_{\hat{t}\hat{t}} = \frac{d\nu(r')}{dr'} e^{-\lambda(r')} = \frac{1}{2f(r)} \frac{df(r)}{dr} \sqrt{f(r)} = \frac{1}{2\sqrt{f(r)}} \frac{df(r)}{dr} \\ &= {}^{14} \frac{\frac{2m}{r^2} - \frac{2}{3}\Lambda r}{2\sqrt{1 - \frac{2m}{r} - \frac{1}{3}\Lambda r^2}} = \frac{3m - \Lambda r^3}{r^2 \sqrt{9 - \frac{18m}{r} - 3\Lambda r^2}} \\ &= \frac{3m - \Lambda r^3}{r^{3/2} \sqrt{9r - 18m - 3\Lambda r^3}} \\ \Gamma^{\hat{r}}_{\hat{\theta}\hat{\theta}} &= \Gamma^{\hat{r}}_{\hat{\phi}\hat{\phi}} = -\Gamma^{\hat{\theta}}_{\hat{r}\hat{\theta}} = -\Gamma^{\hat{\phi}}_{\hat{r}\hat{\phi}} = -\frac{1}{r'} e^{-\lambda(r')} = -\frac{1}{r} \sqrt{f(r)} \\ &= -\frac{1}{r} \sqrt{1 - \frac{2m}{r} - \frac{1}{3}\Lambda r^2} \\ \Gamma^{\hat{\theta}}_{\hat{\phi}\hat{\phi}} &= -\Gamma^{\hat{\phi}}_{\hat{\theta}\hat{\phi}} = -\frac{\cot \theta'}{r'} = -\frac{\cot \theta}{r} \end{aligned}$$

The Ricci tensor

We have $R_{\hat{a}\hat{b}} = \eta_{\hat{a}\hat{b}}\Lambda$ valid in vacuum systems with a cosmological constant, from which we immediately can see that $(\eta_{\hat{a}\hat{b}} = \text{diag}(1, -1, -1, -1))$ $R_{\hat{t}\hat{t}} = -R_{\hat{r}\hat{r}} = -R_{\hat{\theta}\hat{\theta}} = -R_{\hat{\phi}\hat{\phi}} = \Lambda$

11.4.3.2 ¹⁴The general Schwarzschild metric in vacuum with a cosmological constant: The Ricci scalar

The line element

$$ds^2 = e^{2\nu(r)} dt^2 - e^{2\lambda(r)} dr^2 - r^2(d\theta^2 + \sin^2 \theta d\phi^2)$$

In this case we can write the Einstein equation in the local frame (non-coordinate basis):

$$\begin{aligned} 0 &= R_{\hat{a}\hat{b}} - \frac{1}{2}\eta_{\hat{a}\hat{b}}R + \eta_{\hat{a}\hat{b}}\Lambda \\ \Rightarrow 0 &= \eta^{\hat{a}\hat{b}}R_{\hat{a}\hat{b}} - \frac{1}{2}\eta^{\hat{a}\hat{b}}\eta_{\hat{a}\hat{b}}R + \eta^{\hat{a}\hat{b}}\eta_{\hat{a}\hat{b}}\Lambda = R - \frac{1}{2}4R + 4\Lambda \\ \Leftrightarrow R &= 4\Lambda \end{aligned}$$

11.4.3.3 ¹⁵The general Schwarzschild metric in vacuum with a cosmological constant: Integration constants

It can be shown that¹⁵

¹⁴Multiply by $\frac{3r^2}{2}$

¹⁵ From the vacuum equations

$$\begin{aligned} \Rightarrow \quad \lambda(r) &= \ln k - v(r) \\ \lambda'(r) &= -v'(r) \\ e^{-2\lambda(r)} &= \frac{1}{k^2} e^{2v(r)} \end{aligned}$$

We need

$$\begin{aligned} (re^{2v(r)})' &= e^{2v(r)} + 2rv'(r)e^{2v(r)} \\ \Leftrightarrow \quad v'(r) &= \frac{(re^{2v(r)})'}{2re^{2v(r)}} - \frac{1}{2r} \end{aligned}$$

We use the Ricci tensor in the non-coordinate basis for the coordinate $\hat{\theta}$

$$\begin{aligned} R_{\hat{\theta}\hat{\theta}} &= -\frac{v'}{r} e^{-2\lambda(r)} + \frac{\lambda'}{r} e^{-2\lambda(r)} + \frac{(1 - e^{-2\lambda(r)})}{r^2} \\ \Rightarrow \quad \Lambda &= \frac{v'}{r} e^{-2\lambda(r)} - \frac{\lambda'}{r} e^{-2\lambda(r)} - \frac{(1 - e^{-2\lambda(r)})}{r^2} = 2\frac{v'}{k^2 r} e^{2v(r)} - \frac{1}{r^2} + \frac{e^{2v(r)}}{k^2 r^2} \\ &= 2\left(\frac{(re^{2v(r)})'}{2re^{2v(r)}} - \frac{1}{2r}\right) \frac{e^{2v(r)}}{k^2 r} - \frac{1}{r^2} + \frac{e^{2v(r)}}{k^2 r^2} \end{aligned}$$

Renaming

$$\begin{aligned} g(r) &= re^{2v(r)} \\ \Rightarrow \quad \Lambda &= \left(\frac{g'(r)}{g(r)} - \frac{1}{r}\right) \frac{g(r)}{k^2 r^2} - \frac{1}{r^2} + \frac{g(r)}{k^2 r^3} \\ \Leftrightarrow \quad g'(r) &= -k^2 + k^2 \Lambda r^2 \end{aligned}$$

We guess the solution (polynomials with exponents higher than 3 cannot contribute):

$$\begin{aligned} g(r) &= re^{2v(r)} = A + Br + Cr^2 + Dr^3 \\ \Rightarrow \quad g'(r) &= B + 2Cr + 3Dr^2 \\ \Rightarrow \quad -k^2 + k^2 \Lambda r^2 &= B + 2Cr + 3Dr^2 \end{aligned}$$

Now comparing the coefficients we find

$$\begin{aligned} B &= -k^2 \\ C &= 0 \\ D &= \frac{1}{3} k^2 \Lambda \end{aligned}$$

and we can conclude that

$$\begin{aligned} re^{2v(r)} &= A - k^2 r + \frac{1}{3} k^2 \Lambda r^3 \\ \Rightarrow \quad e^{2v(r)} &= \frac{A}{r} - k^2 + \frac{1}{3} k^2 \Lambda r^2 \end{aligned}$$

and the line element

$$ds^2 = e^{2v(r)} dt^2 - e^{2\lambda(r)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

becomes

$$ds^2 = \left(\frac{A}{r} - k^2 + \frac{1}{3} k^2 \Lambda r^2\right) dt^2 - \left(\frac{A}{r} - k^2 + \frac{1}{3} k^2 \Lambda r^2\right)^{-1} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

If $k^2 = -1$ and $\Lambda = 0$ this should be identical to the ordinary Schwarzschild vacuum metric, which means that A has to be equal to: $A = -2m$

11.4.3.4 'The general Schwarzschild metric in vacuum with a cosmological constant: The spatial part of the line element.

The line element

$$ds^2 = e^{2v(r)} dt^2 - e^{2\lambda(r)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

can in Gaussian normal coordinates be written as

$$ds^2 = dt^2 - a^2(t) d\sigma^2$$

In this case we want to find the spatial part of the line element

$$d\sigma^2 = g_{ij} dx_i dx_j$$

and to do that we will use the method^s where the metric is found from the Ricci-tensor:

$$R_{ij} = 2K g_{ij}$$

$$\Leftrightarrow g_{ij} = \frac{1}{2K} R_{ij}$$

$$\Rightarrow g_{rr} = \frac{1}{2K} R_{rr} = \frac{1}{2K} (\Lambda^{\hat{r}}_{\hat{r}})^2 R_{\hat{r}\hat{r}} = -\frac{1}{2K} (e^{\lambda(r)})^2 \Lambda = -\frac{\Lambda}{2K} e^{2\lambda(r)}$$

In another chapter we found that

$$e^{-2\lambda(r)} = \frac{1}{k^2} e^{2\nu(r)} = \frac{1}{k^2} \left(\frac{A}{r} - k^2 + \frac{1}{3} k^2 \Lambda r^2 \right) = \frac{A}{k^2 r} - 1 + \frac{1}{3} \Lambda r^2$$

$$\Rightarrow g_{rr} = -\frac{\Lambda}{2K} \frac{1}{\frac{A}{k^2 r} - 1 + \frac{1}{3} \Lambda r^2}$$

$$g_{\theta\theta} = \frac{1}{2K} R_{\theta\theta} = \frac{1}{2K} (\Lambda^{\hat{\theta}}_{\hat{\theta}})^2 R_{\hat{\theta}\hat{\theta}} = -\frac{1}{2K} (r)^2 \Lambda = -\frac{\Lambda}{2K} r^2$$

$$g_{\phi\phi} = \frac{1}{2K} R_{\phi\phi} = \frac{1}{2K} (\Lambda^{\hat{\phi}}_{\hat{\phi}})^2 R_{\hat{\phi}\hat{\phi}} = -\frac{1}{2K} (r \sin \theta)^2 \Lambda = -\frac{\Lambda}{2K} r^2 \sin^2 \theta$$

$$\Rightarrow d\sigma^2 = -\frac{\Lambda}{2K} \frac{dr^2}{\frac{A}{k^2 r} - 1 + \frac{1}{3} \Lambda r^2} - \frac{\Lambda}{2K} r^2 d\theta^2 - \frac{\Lambda}{2K} r^2 \sin^2 \theta d\phi^2$$

where we can omit the common factor $= -\frac{\Lambda}{2K}$ and finally get if we choose as before $A = -2m$ and $k^2 = -1$

$$d\sigma^2 = \frac{dr^2}{1 - \frac{2m}{r} - \frac{1}{3} \Lambda r^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2$$

11.5 ^tuThe Schwarzschild space-time in Kruskal Coordinates.

11.5.1 The Kruskal coordinates $r > 2m$

$$u = e^{\frac{r}{4m}} \sqrt{\frac{r}{2m} - 1} \cosh \frac{t}{4m}$$

$$v = e^{\frac{r}{4m}} \sqrt{\frac{r}{2m} - 1} \sinh \frac{t}{4m}$$

where

$$u^2 - v^2 = e^{\frac{r}{2m}} \left(\frac{r}{2m} - 1 \right)$$

We calculate

$$\frac{\partial u}{\partial t} = \frac{1}{4m} e^{\frac{r}{4m}} \sqrt{\frac{r}{2m} - 1} \sinh \frac{t}{4m} = \frac{v}{4m}$$

$$\frac{\partial v}{\partial t} = \frac{1}{4m} e^{\frac{r}{4m}} \sqrt{\frac{r}{2m} - 1} \cosh \frac{t}{4m} = \frac{u}{4m}$$

$$\frac{\partial u}{\partial r} = \frac{1}{4m} e^{\frac{r}{4m}} \sqrt{\frac{r}{2m} - 1} \cosh \frac{t}{4m} + e^{\frac{r}{4m}} \frac{1}{2m} \frac{1}{2\sqrt{\frac{r}{2m} - 1}} \cosh \frac{t}{4m} = \frac{u}{4m} \left(\frac{1}{1 - \frac{2m}{r}} \right)$$

$$\frac{\partial v}{\partial r} = \frac{1}{4m} e^{\frac{r}{4m}} \sqrt{\frac{r}{2m} - 1} \sinh \frac{t}{4m} + e^{\frac{r}{4m}} \frac{1}{2m} \frac{1}{2\sqrt{\frac{r}{2m} - 1}} \sinh \frac{t}{4m} = \frac{v}{4m} \left(\frac{1}{1 - \frac{2m}{r}} \right)$$

Now we can use the chain rule

$$du = \frac{\partial u}{\partial t} dt + \frac{\partial u}{\partial r} dr$$

$$dv = \frac{\partial v}{\partial t} dt + \frac{\partial v}{\partial r} dr$$

Written as a matrix

$$\begin{Bmatrix} du \\ dv \end{Bmatrix} = \begin{pmatrix} \frac{\partial u}{\partial t} & \frac{\partial u}{\partial r} \\ \frac{\partial v}{\partial t} & \frac{\partial v}{\partial r} \end{pmatrix} \begin{Bmatrix} dt \\ dr \end{Bmatrix}$$

With the inverse

$$\begin{aligned} \begin{Bmatrix} dt \\ dr \end{Bmatrix} &= {}^{16} \begin{pmatrix} \frac{\partial u}{\partial t} & \frac{\partial u}{\partial r} \\ \frac{\partial v}{\partial t} & \frac{\partial v}{\partial r} \end{pmatrix}^{-1} \begin{Bmatrix} du \\ dv \end{Bmatrix} = \frac{1}{\frac{\partial u}{\partial t} \frac{\partial v}{\partial r} - \frac{\partial u}{\partial r} \frac{\partial v}{\partial t}} \begin{pmatrix} \frac{\partial v}{\partial r} & -\frac{\partial u}{\partial r} \\ -\frac{\partial v}{\partial t} & \frac{\partial u}{\partial t} \end{pmatrix} \begin{Bmatrix} du \\ dv \end{Bmatrix} \\ &= \frac{1}{\left(\frac{v}{4m}\right)^2 \left(\frac{1}{1-\frac{2m}{r}}\right) - \left(\frac{u}{4m}\right)^2 \left(\frac{1}{1-\frac{2m}{r}}\right)} \begin{pmatrix} \frac{v}{4m} \left(\frac{1}{1-\frac{2m}{r}}\right) & -\frac{u}{4m} \left(\frac{1}{1-\frac{2m}{r}}\right) \\ -\frac{u}{4m} & \frac{v}{4m} \end{pmatrix} \begin{Bmatrix} du \\ dv \end{Bmatrix} \\ &= \frac{4m}{v^2 - u^2} \begin{pmatrix} -u \left(1 - \frac{2m}{r}\right) & v \left(1 - \frac{2m}{r}\right) \\ \frac{4m}{v^2 - u^2} & \frac{4m}{v^2 - u^2} \end{pmatrix} \begin{Bmatrix} du \\ dv \end{Bmatrix} \\ \Rightarrow dt &= \frac{4m}{v^2 - u^2} (v du - u dv) \\ \Rightarrow dt^2 &= \frac{16m^2}{(v^2 - u^2)^2} (v^2 du^2 + u^2 dv^2 - 2uv du dv) \\ \Rightarrow dr &= \frac{4m}{v^2 - u^2} \left(1 - \frac{2m}{r}\right) (-u du + v dv) \\ \Rightarrow dr^2 &= \frac{16m^2}{(v^2 - u^2)^2} \left(1 - \frac{2m}{r}\right)^2 (u^2 du^2 + v^2 dv^2 - 2uv du dv) \end{aligned}$$

Next we find

$$\begin{aligned} \left(1 - \frac{2m}{r}\right) dt^2 &= \left(1 - \frac{2m}{r}\right) \frac{16m^2}{(v^2 - u^2)^2} (v^2 du^2 + u^2 dv^2 - 2uv du dv) \\ \left(1 - \frac{2m}{r}\right)^{-1} dr^2 &= \left(1 - \frac{2m}{r}\right) \frac{16m^2}{(v^2 - u^2)^2} (u^2 du^2 + v^2 dv^2 - 2uv du dv) \end{aligned}$$

Substituting into the Schwarzschild line element

$$\begin{aligned} ds^2 &= \left(1 - \frac{2m}{r}\right) dt^2 - \left(1 - \frac{2m}{r}\right)^{-1} dr^2 - r^2 (d\theta + \sin^2 \theta d\phi) \\ &= \left(1 - \frac{2m}{r}\right) \frac{16m^2}{(v^2 - u^2)^2} (v^2 du^2 + u^2 dv^2 - (u^2 du^2 + v^2 dv^2)) - r^2 (d\theta + \sin^2 \theta d\phi) \\ &= \left(1 - \frac{2m}{r}\right) \frac{16m^2}{(u^2 - v^2)} (dv^2 - du^2) - r^2 (d\theta + \sin^2 \theta d\phi) \\ &= 16m^2 \left(1 - \frac{2m}{r}\right) e^{-\frac{r}{2m}} \left(\frac{r}{2m} - 1\right)^{-1} (du^2 - dv^2) - r^2 (d\theta + \sin^2 \theta d\phi) \\ &= \frac{32m^3}{r} e^{-\frac{r}{2m}} (dv^2 - du^2) - r^2 (d\theta + \sin^2 \theta d\phi) \end{aligned}$$

11.5.2 The Kruskal coordinates $r < 2m$

We use the same method as before

¹⁶ $\begin{Bmatrix} a & b \\ c & d \end{Bmatrix}^{-1} = \frac{1}{ad-bc} \begin{Bmatrix} d & -b \\ -c & a \end{Bmatrix}$

$$\begin{aligned}
 u &= e^{\frac{r}{4m}} \sqrt{1 - \frac{r}{2m}} \sinh \frac{t}{4m} \\
 v &= e^{\frac{r}{4m}} \sqrt{1 - \frac{r}{2m}} \cosh \frac{t}{4m} \\
 u^2 - v^2 &= e^{\frac{r}{2m}} \left(\frac{r}{2m} - 1 \right) \\
 \frac{\partial u}{\partial t} &= \frac{1}{4m} e^{\frac{r}{4m}} \sqrt{1 - \frac{r}{2m}} \cosh \frac{t}{4m} = \frac{v}{4m} \\
 \frac{\partial v}{\partial t} &= \frac{1}{4m} e^{\frac{r}{4m}} \sqrt{1 - \frac{r}{2m}} \sinh \frac{t}{4m} = \frac{u}{4m} \\
 \frac{\partial u}{\partial r} &= \frac{1}{4m} e^{\frac{r}{4m}} \sqrt{1 - \frac{r}{2m}} \sinh \frac{t}{4m} - e^{\frac{r}{4m}} \frac{1}{2m} \frac{1}{2\sqrt{1 - \frac{r}{2m}}} \sinh \frac{t}{4m} = \frac{u}{4m} \left(\frac{1}{\frac{2m}{r} - 1} \right) \\
 \frac{\partial v}{\partial r} &= \frac{1}{4m} e^{\frac{r}{4m}} \sqrt{1 - \frac{r}{2m}} \cosh \frac{t}{4m} - e^{\frac{r}{4m}} \frac{1}{2m} \frac{1}{2\sqrt{1 - \frac{r}{2m}}} \cosh \frac{t}{4m} = \frac{v}{4m} \left(\frac{1}{\frac{2m}{r} - 1} \right) \\
 \left\{ \begin{matrix} dt \\ dr \end{matrix} \right\} &= \begin{pmatrix} \frac{\partial u}{\partial t} & \frac{\partial u}{\partial r} \\ \frac{\partial v}{\partial t} & \frac{\partial v}{\partial r} \end{pmatrix}^{-1} \begin{pmatrix} \frac{\partial v}{\partial r} & -\frac{\partial u}{\partial r} \\ -\frac{\partial v}{\partial t} & \frac{\partial u}{\partial t} \end{pmatrix} \begin{pmatrix} du \\ dv \end{pmatrix} \\
 &= \frac{1}{\left(\frac{v}{4m} \right)^2 \left(\frac{1}{\frac{2m}{r} - 1} \right) - \left(\frac{u}{4m} \right)^2 \left(\frac{1}{\frac{2m}{r} - 1} \right)} \begin{pmatrix} \frac{v}{4m} \left(\frac{1}{\frac{2m}{r} - 1} \right) & -\frac{u}{4m} \left(\frac{1}{\frac{2m}{r} - 1} \right) \\ -\frac{u}{4m} & \frac{v}{4m} \end{pmatrix} \begin{pmatrix} du \\ dv \end{pmatrix} \\
 &= \frac{4m}{v^2 - u^2} \begin{pmatrix} -u \left(\frac{2m}{r} - 1 \right) & v \left(\frac{2m}{r} - 1 \right) \end{pmatrix} \begin{pmatrix} du \\ dv \end{pmatrix} \\
 \Rightarrow dt &= \frac{4m}{v^2 - u^2} (vdu - u dv) \\
 \Rightarrow dt^2 &= \frac{16m^2}{(v^2 - u^2)^2} (v^2 du^2 + u^2 dv^2 - 2uvdudv) \\
 \Rightarrow dr &= \frac{4m}{v^2 - u^2} \left(\frac{2m}{r} - 1 \right) (-udu + vdv) \\
 \Rightarrow dr^2 &= \frac{16m^2}{(v^2 - u^2)^2} \left(\frac{2m}{r} - 1 \right)^2 (u^2 du^2 + v^2 dv^2 - 2uvdudv) \\
 \left(1 - \frac{2m}{r} \right) dt^2 &= \left(1 - \frac{2m}{r} \right) \frac{16m^2}{(v^2 - u^2)^2} (v^2 du^2 + u^2 dv^2 - 2uvdudv) \\
 \left(1 - \frac{2m}{r} \right)^{-1} dr^2 &= \left(1 - \frac{2m}{r} \right) \frac{16m^2}{(v^2 - u^2)^2} (u^2 du^2 + v^2 dv^2 - 2uvdudv) \\
 \Rightarrow ds^2 &= \left(1 - \frac{2m}{r} \right) dt^2 - \left(1 - \frac{2m}{r} \right)^{-1} dr^2 - r^2 (d\theta + \sin^2 \theta d\phi) \\
 &= \frac{32m^3}{r} e^{-\frac{r}{2m}} (dv^2 - du^2) - r^2 (d\theta + \sin^2 \theta d\phi)
 \end{aligned}$$

11.6 *Eddington-Finkelstein Coordinates – Avoiding the singularity

In order to avoid problems in the $r = 2m$ singularity in the Schwarzschild space-time we can use a transformation of coordinates: The Eddington-Finkelstein coordinates, which introduces a new coordinate $v(t, r)$ defined by

$$t = v - r - 2m \ln \left| \frac{r}{2m} - 1 \right|$$

Where the numerical value ensures that the transformation is valid both outside and inside the horizon $r = 2m$. Also notice that $t \rightarrow \infty$ when $r \rightarrow 2m$ which we saw in the Schwarzschild solution.

11.6.1 *The Line-element

$$t = v - r - 2m \ln \left| \frac{r}{2m} - 1 \right|$$

$$\Rightarrow dt = dv - dr - d \left(2m \ln \left| \frac{r}{2m} - 1 \right| \right) = dv - dr - \frac{1}{\frac{r}{2m} - 1} dr = dv - \frac{1}{1 - \frac{2m}{r}} dr$$

$$\Rightarrow dt^2 = \left(dv - \frac{1}{1 - \frac{2m}{r}} dr \right)^2 = dv^2 + \left(\frac{1}{1 - \frac{2m}{r}} \right)^2 dr^2 - 2 \frac{1}{1 - \frac{2m}{r}} dv dr$$

Substituting this into the Schwarzschild line-element

$$ds^2 = - \left(1 - \frac{2m}{r} \right) dt^2 + \left(1 - \frac{2m}{r} \right)^{-1} dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\phi^2)$$

$$= {}^{17} {}^{18} - \left(1 - \frac{2m}{r} \right) dv^2 + 2dvdr + r^2 (d\theta^2 + \sin^2 \theta d\phi^2)$$

11.6.2 *The Radial Null Geodesics

The radial null geodesics implies that $ds^2 = d\theta = d\phi = 0$

$$\Rightarrow 0 = - \left(1 - \frac{2m}{r} \right) dv^2 + 2dvdr = \left(- \left(1 - \frac{2m}{r} \right) dv + 2dr \right) dv$$

$dv = 0$:

This solution leads to the familiar ingoing null rays in the Schwarzschild solution

$$dt^2 = \left(\frac{1}{1 - \frac{2m}{r}} \right)^2 dr^2$$

$r = 2m$:

$$\Rightarrow \frac{dv}{dr} = 0$$

which are light rays that are neither ingoing nor outgoing.

Next we solve the differential equation:

$$0 = - \left(1 - \frac{2m}{r} \right) dv + 2dr$$

$$\Rightarrow \frac{dv}{dr} = 2 \left(1 - \frac{2m}{r} \right)^{-1} = \frac{2r}{r - 2m}$$

Notice if $r > 2m$ then $\frac{dv}{dr} > 0$ and the radial light rays are outgoing

Notice if $r < 2m$ then $\frac{dv}{dr} < 0$ and the radial light rays are ingoing

$${}^{17} = - \left(1 - \frac{2m}{r} \right) \left(dv^2 + \left(\frac{1}{1 - \frac{2m}{r}} \right)^2 dr^2 - 2 \frac{1}{1 - \frac{2m}{r}} dv dr \right) + \left(1 - \frac{2m}{r} \right)^{-1} dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\phi^2) =$$

¹⁸ Notice: For $r \rightarrow \infty$ the Eddington-Finkelstein space-time approaches the flat space-time: $ds^2 = -dv^2 + 2dvdr + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi$ which is the same as flat Minkowsky space-time.

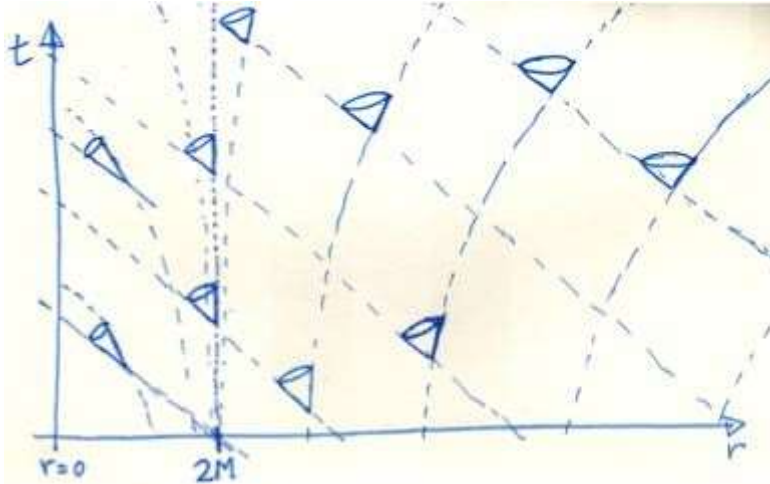
Solving

$$dv = \frac{2r}{r-2m} dr$$

$$\Rightarrow v - v_0 \stackrel{19}{=} 2(r + 2m \ln|r - 2m|)$$

In summary: For $r > 2m$ we have both ingoing and outgoing radial light rays, the ingoing light rays can either pass the event horizon or stop at the event horizon. The event horizon is so to speak a one way membrane and no light can go out again. If $r < 2m$ all the light rays are ingoing.

As illustrated by the drawing below^y the radial null-geodesic points both outward and inward outside the event horizon but only inward inside the event horizon. Also notice that for an astronaut his trajectory always lies inside the light cone so the same rules applies for him.



11.7 The Kerr metric

11.7.1 The Kerr-Newman geometry

A more general metric is the Kerr-Newman geometry, corresponding to a simultaneously rotating and electrically charged black hole of mass m , charge Q and angular momentum S .

$$ds^2 = \frac{\Delta}{\Sigma} (dt - a \sin^2 \theta d\phi)^2 - \frac{\sin^2 \theta}{\Sigma} ((r^2 + a^2)d\phi - a dt)^2 - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2$$

$$= \left(1 - \frac{2mr}{\Sigma} + \frac{Q^2}{\Sigma}\right) dt^2 + \frac{4amr \sin^2 \theta}{\Sigma} dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 - \left(r^2 + a^2 + \frac{2a^2 m r \sin^2 \theta}{\Sigma}\right) \sin^2 \theta d\phi^2$$

$$\Delta = r^2 - 2mr + a^2 + Q^2$$

¹⁹ (Spiegel, 1990) (14.60) $\int \frac{xdx}{ax+b} = \frac{x}{a} - \frac{b}{a^2} \ln(ax+b)$

²⁰ $= \frac{\Delta}{\Sigma} [dt^2 + (a \sin^2 \theta)^2 d\phi^2 - 2a \sin^2 \theta dt d\phi] - \frac{\sin^2 \theta}{\Sigma} [a^2 dt^2 + (r^2 + a^2)^2 d\phi^2 - 2a(r^2 + a^2) dt d\phi] - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 = \frac{1}{\Sigma} [\Delta - a^2 \sin^2 \theta] dt^2 + \frac{1}{\Sigma} [-\Delta 2a \sin^2 \theta + 2 \sin^2 \theta a(r^2 + a^2)] dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 + \frac{1}{\Sigma} [\Delta (a \sin^2 \theta)^2 - \sin^2 \theta (r^2 + a^2)^2] d\phi^2 =$

²¹ $= \frac{1}{\Sigma} [\Delta - a^2 \sin^2 \theta] dt^2 + \frac{2a \sin^2 \theta}{\Sigma} [-\Delta + (r^2 + a^2)] dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 + \frac{\sin^2 \theta}{\Sigma} [\Delta a^2 \sin^2 \theta - (r^2 + a^2)^2] d\phi^2 =$

²² $= \frac{1}{\Sigma} [(\Sigma - 2mr + a^2 \sin^2 \theta + Q^2) - a^2 \sin^2 \theta] dt^2 + \frac{2a \sin^2 \theta}{\Sigma} [-(\Sigma - 2mr + a^2 \sin^2 \theta) + (\Sigma + a^2 \sin^2 \theta)] dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 + \frac{\sin^2 \theta}{\Sigma} [(\Sigma - 2mr + a^2 \sin^2 \theta) a^2 \sin^2 \theta - (\Sigma + a^2 \sin^2 \theta)^2] d\phi^2 =$

²³ $= \left(1 - \frac{2mr}{\Sigma} + \frac{Q^2}{\Sigma}\right) dt^2 + \frac{4amr \sin^2 \theta}{\Sigma} dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 + \frac{\sin^2 \theta}{\Sigma} [\Sigma a^2 \sin^2 \theta - 2mra^2 \sin^2 \theta + (a^2 \sin^2 \theta)^2 - (\Sigma^2 + (a^2 \sin^2 \theta)^2 + 2\Sigma a^2 \sin^2 \theta)] d\phi^2$

²⁴ $= \left(1 - \frac{2mr}{\Sigma} + \frac{Q^2}{\Sigma}\right) dt^2 + \frac{4amr \sin^2 \theta}{\Sigma} dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 + \frac{\sin^2 \theta}{\Sigma} [-\Sigma a^2 \sin^2 \theta - 2mra^2 \sin^2 \theta - \Sigma^2] d\phi^2 =$

²⁵ $= \left(1 - \frac{2mr}{\Sigma} + \frac{Q^2}{\Sigma}\right) dt^2 + \frac{4amr \sin^2 \theta}{\Sigma} dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 - \sin^2 \theta \left[a^2 \sin^2 \theta + \frac{2mra^2 \sin^2 \theta}{\Sigma} + \Sigma \right] d\phi^2 =$

$$\begin{aligned}\Sigma &= r^2 + a^2 \cos^2 \theta = r^2 + a^2 - a^2 \sin^2 \theta = \Delta - Q^2 + 2mr - a^2 \sin^2 \theta \\ a &= \frac{S}{m}\end{aligned}$$

11.7.1.1 $Q = 0$

In the case of $Q = 0$ we see immediately that the Kerr-Newman geometry reduces to the Kerr geometry describing a non-charged rotating black hole.

$$\begin{aligned}ds^2 &= \left(1 - \frac{2mr}{\Sigma}\right) dt^2 + \frac{4amr \sin^2 \theta}{\Sigma} dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 - \left(r^2 + a^2 + \frac{2a^2mr \sin^2 \theta}{\Sigma}\right) \sin^2 \theta d\phi^2 \\ \Delta &= r^2 - 2mr + a^2 \\ \Sigma &= r^2 + a^2 \cos^2 \theta = r^2 + a^2 - a^2 \sin^2 \theta = \Delta + 2mr - a^2 \sin^2 \theta \\ a &= \frac{S}{m}\end{aligned}$$

11.7.1.2 $S = 0$

In the case of $S = 0$ we see immediately that the Kerr-Newman geometry reduces to the Reissner-Nordström geometry describing a charged non-rotating black hole.

$$\begin{aligned}ds^2 &= {}^{26} \left(1 - \frac{2mr}{\Sigma} + \frac{Q^2}{\Sigma}\right) dt^2 - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 - r^2 \sin^2 \theta d\phi^2 \\ &= \left(1 - \frac{2m}{r} + \frac{Q^2}{r^2}\right) dt^2 - \frac{1}{1 - \frac{2m}{r} + \frac{Q^2}{r^2}} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2 \\ \Delta &= r^2 - 2mr + Q^2 \\ \Sigma &= r^2\end{aligned}$$

11.7.1.3 $Q = 0$ and $S = 0$

In the case of $Q = 0$ and $S = 0$ we see immediately that the Kerr-Newman geometry reduces to the Schwarzschild geometry describing a non-charged non-rotating black hole.

$$\begin{aligned}ds^2 &= {}^{27} \left(1 - \frac{2m}{r}\right) dt^2 - \frac{1}{1 - \frac{2m}{r}} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2 \\ \Delta &= r^2 - 2mr \\ \Sigma &= r^2\end{aligned}$$

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²⁶ = $\left(1 - \frac{2mr}{\Sigma} + \frac{Q^2}{\Sigma}\right) dt^2 + \frac{4amr \sin^2 \theta}{\Sigma} dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 - \left(r^2 + a^2 + \frac{2a^2mr \sin^2 \theta}{\Sigma}\right) \sin^2 \theta d\phi^2 =$

²⁷ = $\left(1 - \frac{2m}{r}\right) dt^2 + \frac{4amr \sin^2 \theta}{\Sigma} dt d\phi - \frac{\Sigma}{\Delta} dr^2 - \Sigma d\theta^2 - \left(r^2 + a^2 + \frac{2a^2mr \sin^2 \theta}{\Sigma}\right) \sin^2 \theta d\phi^2 =$

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- ^a (A.S.Eddington, 1924, p. 83)
^b (McMahon, 2006, s. 231)
^c (McMahon, 2006, s. 204)
^d (McMahon, 2006, s. 204)
^e (McMahon, 2006, s. 211)
^f (McMahon, 2006, s. 218)
^g (Hartle, 2003, p. 187)
^h (Hartle, 2003, p. 166)
ⁱ (McMahon, 2006, s. 216)
^j (McMahon, 2006, s. 215), (Hartle, 2003, p. 546)
^k (McMahon, 2006, s. 231), (Hartle, 2003, s. 546)
^l (d'Inverno, 1992, p. 90)
^m (McMahon, 2006, s. 216), equation (10.35)
ⁿ (Hartle, 2003, p. 183)
^o (McMahon, 2006, s. 231-32)
^p (McMahon, 2006, s. 277)
^q (McMahon, 2006, s. 277)
^r (McMahon, 2006, s. 277)
^s (McMahon, 2006, s. 260)
^t (McMahon, 2006, s. 242)
^u (Hartle, An Introduction to Einstein's General Relativity, 2003, s. 279)
^v (McMahon, 2006, s. 239), (Hartle, An Introduction to Einstein's General Relativity, 2003, s. 256), (Carroll, 2004, p. 221)
^w (Hartle, An Introduction to Einstein's General Relativity, 2003, s. 277)
^x (Hartle, An Introduction to Einstein's General Relativity, 2003, s. 259)
^y <https://physics.stackexchange.com/questions/57726/future-light-cones-inside-black-hole>
^z (C.W.Misner, 1973) chapter 33